Search for Neutrino-Induced Neutral-Current $\Delta$ Radiative Decay in MicroBooNE and a First Test of the MiniBooNE Low Energy Excess under a Single-Photon Hypothesis


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We report results from a search for neutrino-induced neutral current (NC) resonant $\Delta$(1232) baryon production followed by $\Delta$ radiative decay, with a (0.8) GeV neutrino beam. Data corresponding to MicroBooNE’s first three years of operations ($6.80 \times 10^{20}$ protons on target) are used to select single-photon events with one or zero protons and without charged leptons in the final state (1$\gamma1p$ and 1$0p$, respectively). The background is constrained via an in situ high-purity measurement of NC $\pi^0$ events, made possible via dedicated 2$\gamma1p$ and 2$0p$ selections. A total of 16 and 153 events are observed for the 1$\gamma1p$ and 1$0p$ selections, respectively, compared to a constrained background prediction of 20.5 $\pm$ 3.65(syst) and 145.1 $\pm$ 13.8(syst) events. The data lead to a bound on an anomalous enhancement of the normalization of NC $\Delta$ radiative decay of less than 2.3 times the predicted nominal rate for this process at the 90% confidence level (C.L.). The measurement disfavors a candidate photon interpretation of the MiniBooNE low-energy excess as a factor of 3.18 times the nominal NC $\Delta$ radiative decay rate at the 94.8% C.L., in favor of the nominal prediction, and represents a greater than 50-fold improvement over the world’s best limit on single-photon production in NC interactions in the sub-GeV neutrino energy range.

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For over two decades, the anomalous signals consisting of MiniBooNE’s low-energy excess (LEE) [1–3] and the prior LSND [4] $\bar{\nu}_e$ appearance results have been at the forefront of neutrino physics. Each has been interpreted as evidence for new types of neutrinos or other physics beyond the standard model (SM). The existence of new particles would be the first evidence for a new paradigm of physics associated with the neutrino sector since the discovery of neutrinos mass via their observed oscillations, and would have profound ramifications for all particle physics, astrophysics, and cosmology. At the heart of this puzzle of anomalies in need of interpretation is the fact MiniBooNE could not differentiate neutrino interactions producing an electron (such as from $\nu_e$ appearance due to light sterile neutrinos) from those with a single photon in the final state. Thus, both types of interactions must be examined independently as a source of the LEE.

Neutrino-induced neutral current (NC) production of the $\Delta$(1232) baryon resonance with subsequent $\Delta$ radiative decay is predicted to be the dominant source of single photons in neutrino-argon scattering below 1 GeV [5]. Although $\Delta$ radiative decay is predicted in the SM, and measurements of photoproduction [6] and virtual compton scattering [7] are well described by theory, this process has never been directly observed in neutrino scattering. Previous searches have been performed by the T2K [8] and NOMAD [9] experiments with average incident neutrino energies $E_\nu$ of 0.85 and 25 GeV, respectively, resulting in leading limits on this process. Although on a different target, T2K’s result is closest in $E_\nu$ to that of the MiniBooNE beam. However, the 90% confidence level (C.L.) limit is $\sim100$ times the theoretically predicted rate of NC $\Delta$ radiative decay.

In this Letter, we present the world’s most sensitive search for NC $\Delta \rightarrow N\gamma$, where $N = p, n$, using neutrino-argon scattering data collected by the MicroBooNE detector [10]. MicroBooNE is an 85 metric ton active volume liquid argon time projection chamber (LArTPC) situated at a similar baseline in the same muon neutrino dominated Booster Neutrino Beam (BNB) at Fermilab [11] as MiniBooNE, with $\langle E_\nu \rangle = 0.8$ GeV. The measurement
makes use of data corresponding to a BNB exposure of $6.80 \times 10^{30}$ protons on target (POT), collected during 2016–2018. LArTPC technology allows MicroBooNE to distinguish electromagnetic showers originating from electrons or photons based on ionization energy deposition ($dE/dx$) at the start of the shower, and the nonzero conversion distance of the photon relative to the interaction vertex.

This search represents a first for this process with argon as the neutrino target, and also constitutes the first test of the MiniBooNE LEE under a single-photon interpretation. In a fit to the radial distribution of the MiniBooNE data with statistical errors only, an enhancement of NC $\Delta \rightarrow N\gamma$ (as predicted by the NUANCE [12] neutrino event generator on CH$_2$) by a normalization factor of $x_{MB} = 3.18$ (quoted with no uncertainty) was found to provide the best fit for the observed LEE [3]. We perform an explicitly model-dependent test of this interpretation, cast as a factor of 3.18 enhancement to the predicted NC $\Delta \rightarrow N\gamma$ rate in MicroBooNE, under a two-hypothesis $\chi^2$ test between the enhanced rate and the nominal NC $\Delta \rightarrow N\gamma$ prediction.

MicroBooNE uses a custom tune [13] of the GENIE neutrino event generator v3.0.6 [14,15] to simulate neutrino-argon interactions. At BNB energies, the dominant source of single-photon production with no charged leptons or pions in the final state is NC $\Delta(1232) \rightarrow N\gamma$. This process is included in the MicroBooNE nominal prediction exactly as modeled in GENIE. Heavier resonances and nonresonant processes, including coherent single-photon production [16], are not currently included in the simulation, but are each estimated to contribute at the 10% level or less. Both these processes would produce slightly higher-energy photons than the $\Delta(1232)$ resonance, and a more forward (in the direction of the neutrino beam) photon in the case of coherent production. Although such events may be selected by this analysis, we do not explicitly quantify their selection efficiency and in this Letter we focus on the dominant NC $\Delta(1232) \rightarrow N\gamma$ process.

The MicroBooNE NC $\Delta \rightarrow N\gamma$ search exclusively targets events with a single, photonlike electromagnetic shower and either no other visible activity or one visible final-state proton. These are referred to as $1\gamma 0p$ and $1\gamma 1p$ events and primarily probe $\Delta \rightarrow n\gamma$ and $\Delta \rightarrow p\gamma$ decays, respectively. The analysis selects and simultaneously fits $1\gamma 1p$ and $1\gamma 0p$ data-to-Monte Carlo (MC) simulated distributions together with two additional, mutually exclusive but highly correlated event samples: $2\gamma 1p$, and $2\gamma 0p$. The signal, defined as all true NC $\Delta \rightarrow N\gamma$ events whose true interaction vertex is inside the active TPC, contributes predominantly to the $1\gamma$ event samples. The high-statistics $2\gamma$ samples are enhanced in NC $\Delta \rightarrow N\pi^0$ production, which is the dominant source of misidentified background to the $1\gamma 1p$ and $1\gamma 0p$ event samples.

Reconstruction of all four event samples makes use of the Pandora framework [17]. Reconstructed ionization charge hits are clustered and matched across three 2D projected views of the MicroBooNE active TPC volume into 3D reconstructed objects. These are then classified as tracks or showers based on a multivariate classifier score and aggregated into candidate neutrino interactions. The topological selection of interactions with exactly one shower and zero or one tracks represents the basis of the $1\gamma$ selections. Subsequently, preselection requires that the reconstructed vertex, shower-start point and track (as applicable) are all fully contained within the detector fiducial volume. A minimum energy requirement is imposed on the shower, ensuring good reconstruction performance, and a maximum track length requirement is imposed on the track, rejecting obvious muon backgrounds. Tracks are also required to have a high $dE/dx$ consistent with that of a proton. Finally, an opening angle requirement between the track and shower directions is applied to eliminate colinear events where the start of a shower can be misreconstructed as a track.

The preselected events are fed into a set of boosted decision trees (BDTs), each designed to reject a distinct background and select NC $\Delta \rightarrow N\gamma$ events. The gradient boosting algorithm XGBoost [18] is used to train the BDTs. A cosmic BDT rejects cosmogenic backgrounds and is trained on cosmic ray data events collected when no neutrino beam was present. Track calorimetry is used to reject cosmic muons, with track and shower directionality-based variables proving powerful discriminators. A NC $\pi^0$ BDT compares the relationship of the reconstructed shower and track to those expected from $\pi^0$ decay kinematics to separate true single-photon events from those containing a $\pi^0$ decay where a second photon is not reconstructed. A charged current (CC) $\nu_e$ BDT targets the intrinsic $\nu_e$ background events. Here, the photon conversion distance and shower calorimetry play important roles. A fourth BDT is designed to veto events in which a second shower from a $\pi^0$ decay deposits some charge, but fails 3D shower reconstruction. Such events can result in 2D charge hits near the neutrino interaction that are not associated with a 3D object. A plane-by-plane clustering algorithm, DBSCAN [19], is used to group these unassociated hits, and properties including direction, shape, and energy of the cluster are used to determine consistency with a second shower from a $\pi^0$ decay. A final CC $\nu_\mu$-focused BDT removes any remaining backgrounds, primarily targeting the muon track through track calorimetry variables.

The $1\gamma 1p$ selection uses all five BDTs. The absence of a track in the $1\gamma 0p$ sample means that the $1\gamma 0p$ selection cannot use these BDTs identically, as it is limited to only shower variables. As such, it uses variations of the cosmic and NC$\pi^0$ BDTs, and a third BDT merging the functionality of the CC $\nu_e$ and CC $\nu_\mu$-focused BDTs, targeting all remaining backgrounds. All BDTs are trained explicitly to select well-reconstructed NC $\Delta \rightarrow N\gamma$ events. While model dependent, this leverages the kinematics and correlations
The expected event rates in the $1\gamma p$ and $1\gamma 0p$ samples. “Dirt (outside TPC)” represents any neutrino interaction that originates outside the active TPC, but scatters inside. Relative to the topological selection stage, the $\nu_\mu$ CC rejection is 99.8\% and 87.6\% for $1\gamma p$ and $1\gamma 0p$, respectively.

<table>
<thead>
<tr>
<th>Process</th>
<th>$1\gamma p$</th>
<th>$1\gamma 0p$</th>
</tr>
</thead>
<tbody>
<tr>
<td>NC $1\pi^0$ Noncoherent</td>
<td>24.0</td>
<td>68.1</td>
</tr>
<tr>
<td>NC $1\pi^0$ Coherent</td>
<td>0.0</td>
<td>7.6</td>
</tr>
<tr>
<td>CC $\nu_\mu$ $1\pi^0$</td>
<td>0.5</td>
<td>14.0</td>
</tr>
<tr>
<td>CC $\nu_\epsilon$ and $\bar{\nu}_\epsilon$</td>
<td>0.4</td>
<td>11.1</td>
</tr>
<tr>
<td>BNB other</td>
<td>2.1</td>
<td>18.1</td>
</tr>
<tr>
<td>Dirt (outside TPC)</td>
<td>0.0</td>
<td>36.4</td>
</tr>
<tr>
<td>Cosmic ray data</td>
<td>0.0</td>
<td>10.0</td>
</tr>
<tr>
<td>Total background (unconstr.)</td>
<td>27.0</td>
<td>165.4</td>
</tr>
<tr>
<td>NC $\Delta \to N\gamma$</td>
<td>4.88</td>
<td>6.55</td>
</tr>
</tbody>
</table>
The one-bin $1\gamma 1p$ and $1\gamma 0p$ event rates are fit simultaneously with the $2\gamma 1p$ and $2\gamma 0p$ distributions shown in Fig. 1. The fit makes use of a covariance matrix that encapsulates statistical and systematic uncertainties and bin-to-bin correlations, allowing for both the expected rate and uncertainties of the NC $\pi^0$ backgrounds in the $1\gamma$ samples to be effectively constrained by the high-statistics data observed in the $2\gamma$ samples.

The normalization $x_\Delta$ can also be reinterpreted in various ways. First, it can be reinterpreted as a scaling of an effective branching fraction $B_{\text{eff}}(\Delta \rightarrow N\gamma)$, where the nominal prediction ($x_\Delta = 1$) corresponds to an effective branching fraction of 0.6%. This effective branching fraction can be thought of as a metric to account for any uncertain nuclear effects that might modify the $\Delta$ behavior inside the nuclear medium, as we cannot observe the true $\Delta \rightarrow N\gamma$ branching fraction directly. In addition, any BSM effect that can contribute as an NC $\Delta$-like process (with a single photonlike shower in the final state) could lead to an effective modification to the observed branching fraction. Although GENIE prescribes a normalization uncertainty for $B_{\text{eff}}(\Delta \rightarrow N\gamma)$, this uncertainty is not included in the fit. In addition, with the knowledge that GENIE predicts a cross section for NC $\Delta \rightarrow N\gamma$ production to be $\sigma_{\text{NC}\Delta-N\gamma}^{\text{GENIE}} = 8.61 \times 10^{-42} \text{ cm}^{-2}/\text{nucleon}$, we can also reinterpret $x_\Delta$ as scaling on this production cross section. The Feldman-Cousins [21] approach is followed to construct the confidence intervals for $x_\Delta$ given the best fit to the observed data, with a metric of $\Delta \chi^2$ defined using the combined-Neyman-Pearson $\chi^2$ [22] as an approximation of the log-likelihood ratio.

Systematic uncertainties include contributions from flux, cross-section modeling, hadron reinteractions, detector effects, and finite statistics used in the background predictions (both MC and cosmic ray data). The flux uncertainties incorporate hadron production uncertainties, uncertainties on pion and nucleon scattering in the beryllium target and surrounding aluminum magnetic horn, and mis-modeling of the horn current. Following Ref. [23], these are implemented by reweighting the flux prediction $\sigma_{\text{NC}1p}$ and $\sigma_{\text{NC}0p}$ total background events are summarized in Table III. The GENIE cross-section uncertainties dominate. This stems from the uncertainties on NC $\pi^0$ production on argon, which forms the largest background and has not been measured to high precision to date. Both cross-section and flux uncertainties are strongly correlated between the $1\gamma$ and $2\gamma$ event samples. The simultaneous fit to the $1\gamma$ and $2\gamma$ samples is equivalent to a $1\gamma$-only fit where the background and uncertainty are conditionally constrained [29] by the $2\gamma$ samples. Given the $2\gamma$ samples’ statistics, this constraint effectively reduces the total background systematic uncertainty of the $1\gamma 1p$ and $1\gamma 0p$ samples by 40% and 50%, and the total background prediction by 24.1% and 12.3%, respectively.

The 90% C.L. sensitivity is quantified for a Feldman-Cousins-corrected limit in the case of a background-only observation, $x_\Delta = 0$, to be less than $x_\Delta = 2.5$, corresponding to $B_{\text{eff}}(\Delta \rightarrow N\gamma) = 1.50\%$ and $\sigma_{\text{NC}\Delta-N\gamma}^{\text{MC}} = 21.5 \times 10^{-42} \text{ cm}^{-2}/\text{nucleon}$. Under a two-hypothesis $\Delta \chi^2$ test, the expected sensitivity of the median experiment assuming the nominal prediction, to reject the LEE hypothesis ($x_{\text{MB}} = 3.18$) in favor of the nominal hypothesis ($x_{\Delta} = 1$) is $1.5\sigma$; in the case of the median experiment assuming the LEE hypothesis, the sensitivity to reject the nominal hypothesis in favor of the LEE hypothesis is $1.6\sigma$.

The reconstruction, selection, and fitting methods employed in this search were developed adhering to a signal-blind analysis strategy, whereby the data were kept blind until the analysis was fully developed, with the

<table>
<thead>
<tr>
<th>Type of uncertainty</th>
<th>$1\gamma 1p$</th>
<th>$1\gamma 0p$</th>
</tr>
</thead>
<tbody>
<tr>
<td>Flux model</td>
<td>7.4%</td>
<td>6.6%</td>
</tr>
<tr>
<td>GENIE cross-section model</td>
<td>24.8%</td>
<td>16.3%</td>
</tr>
<tr>
<td>GEANT4 reinteractions</td>
<td>1.1%</td>
<td>1.3%</td>
</tr>
<tr>
<td>Detector effects</td>
<td>12.2%</td>
<td>6.4%</td>
</tr>
<tr>
<td>Finite background statistics</td>
<td>8.3%</td>
<td>4.0%</td>
</tr>
<tr>
<td>Total uncertainty (unconstr.)</td>
<td>29.8%</td>
<td>19.2%</td>
</tr>
<tr>
<td>Total uncertainty (constr.)</td>
<td>17.8%</td>
<td>9.5%</td>
</tr>
</tbody>
</table>
 exception of a small subset of the data consisting of $0.51 \times 10^{20}$ POT, used for analysis validation. After $1\gamma p$ and $1\gamma 0p$ event samples were unblinded, 16 data events with an expected constrained background of $20.5 \pm 3.6$ (syst) events were observed in the $1\gamma p$ event sample, and 153 data events with an expected constrained background of $145.1 \pm 13.8$ (syst) events were observed in the $1\gamma 0p$ event sample. The reconstructed shower energy distributions of selected $1\gamma p$ and $1\gamma 0p$ events are shown in Fig. 2. Overall, a systematic deficit of data relative to the unconstrained MC prediction is observed, which is within systematic and statistical uncertainties, and consistent with a similar deficit in the $2\gamma$ event samples. The expected signal and background predictions are summarized in Table IV and Fig. 3, and compared to the observed data, both before and after applying the $2\gamma$ conditional constraint. The $2\gamma$ constraint reduces the total background prediction, consistently with

![FIG. 2. Energy spectra for the (a) $1\gamma p$ and (b) $1\gamma 0p$ selected events. The upper section in each figure shows the unconstrained background predictions and breakdowns as a function of reconstructed shower energy. The lower section shows the total background predictions and breakdowns as a function of reconstructed shower energy.](image)

![FIG. 3. The observed event rates for the (a) $1\gamma p$ and (b) $1\gamma 0p$ event samples, and comparisons to unconstrained (left) and constrained (right) background and LEE model predictions. The event rates are the sum of all events with reconstructed shower energy between 0–600 MeV and 100–700 MeV for (a) and (b), respectively. The one-bin background only conditionally constrained $\chi^2$ is 0.63 and 0.18 for $1\gamma p$ and $1\gamma 0p$, respectively.](image)
FIG. 4. The resulting $\Delta \chi^2$ curve after fitting to $x_\Delta$ using the Feldman-Cousins procedure, showing extracted confidence intervals. The best fit is found to be at $x_\Delta = 0$ with a $\chi^2_{bf} = 5.53$. Shown also is the reinterpretation of this scaling factor as both an effective branching fraction, $B_{\text{eff}}(\Delta \rightarrow N\gamma)$, and a cross section, $\sigma_{N\text{C} \rightarrow N\gamma}^{\text{eff}}$. The default GENIE value corresponds to $x_\Delta = 1$. The error on the LEE model is estimated from the MiniBooNE result [3] with statistical and systematic uncertainty. It should be noted that this uncertainty does not account for systematic correlations between MiniBooNE and MicroBooNE.

The data to MC simulation ratio observed in the 2$\gamma$ event samples.

The best-fit value for $x_\Delta$ obtained from the fit is 0, with a $\chi^2_{bf}$ of 5.53 for 15 degrees of freedom (d.o.f.). This measurement is in agreement with the nominal NC $\Delta \rightarrow N\gamma$ rate (corresponding to $x_\Delta = 1$) within 1$\sigma$ (67.8% C.L.) with a $\chi^2$ of 6.47 for 16 d.o.f. The Feldman-Cousins calculated confidence limit leads to a one-sided bound on the normalization of NC $\Delta \rightarrow N\gamma$ events of $x_\Delta < 2.3$, corresponding to $B_{\text{eff}}(\Delta \rightarrow N\gamma) < 1.38\%$ and $\sigma_{NC \rightarrow N\gamma}^{\text{eff}} < 19.8 \times 10^{-42}$ cm$^2$/nucleon, at 90% C.L. This is summarized in Fig. 4.

This result represents the most stringent limit on neutrino-induced NC $\Delta \rightarrow N\gamma$ on any nuclear target [8,9], and a significant improvement over previous searches, in particular in the neutrino energy range below 1 GeV. Under a two-hypothesis test, the data rule out the interpretation of the MiniBooNE anomalous excess [30] as a factor of 3.18 enhancement to the rate of $\Delta \rightarrow N\gamma$, in favor of the nominal prediction at 94.8% C.L. (1.9$\sigma$). While this is a model-dependent test of the MiniBooNE LEE, and does not apply universally to all other photonlike interpretations, it provides an important constraint on this process and a first direct test of the MiniBooNE LEE, and opens the door to further searches that focus on a broader range of models. Those include coherent single-photon production [5], anomalous contributions of which could give rise to additional events and would be expected to leave an imprint in the 1$\gamma$0$\pi$ selection, as well as more exotic Beyond-SM processes that manifest as single-photon events, such as colinear $e^+e^-$ pairs from $Z'$ [31,32] or scalar [33] decays, among others. Follow-up MicroBooNE analyses will explicitly target and quantify sensitivity to these alternative hypotheses, as well as model-independent single-photon searches.

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[20] See Supplemental Material at http://link.aps.org/supplemental/10.1103/PhysRevLett.128.111801 for a description of the two shower NC\(\pi^0\) rich selections as well as a data release corresponding to the results shown in this paper.


