

Original citation:

Atlas Collaboration (Including: Farrington, Sinead, Harrison, Paul F., Janus, M., Jeske, C., Jones, G. (Graham), Martin, T. A. and Pianori, E.). (2013) Search for charginos nearly mass degenerate with the lightest neutralino based on a disappearing-track signature in pp collisions at $(\sqrt{s})=8$ TeV with the ATLAS detector. European Physical Journal C, Volume 88 (Number 11). Article number 112006.

Permanent WRAP url:

<http://wrap.warwick.ac.uk/58980>

Copyright and reuse:

The Warwick Research Archive Portal (WRAP) makes this work of researchers of the University of Warwick available open access under the following conditions.

This article is made available under the Creative Commons Attribution- 3.0 Unported (CC BY 3.0) license and may be reused according to the conditions of the license. For more details see <http://creativecommons.org/licenses/by/3.0/>

A note on versions:

The version presented in WRAP is the published version, or, version of record, and may be cited as it appears here.

For more information, please contact the WRAP Team at: publications@warwick.ac.uk

warwick**publications**wrap

highlight your research

<http://wrap.warwick.ac.uk/>

Search for charginos nearly mass degenerate with the lightest neutralino based on a disappearing-track signature in pp collisions at $\sqrt{s} = 8$ TeV with the ATLAS detector

G. Aad *et al.**

(ATLAS Collaboration)

(Received 14 October 2013; published 6 December 2013)

A search is presented for direct chargino production based on a disappearing-track signature using 20.3 fb^{-1} of proton-proton collisions at $\sqrt{s} = 8$ TeV collected with the ATLAS experiment at the LHC. In anomaly-mediated supersymmetry breaking (AMSB) models, the lightest chargino is nearly mass degenerate with the lightest neutralino and its lifetime is long enough to be detected in the tracking detectors by identifying decays that result in tracks with no associated hits in the outer region of the tracking system. Some models with supersymmetry also predict charginos with a significant lifetime. This analysis attains sensitivity for charginos with a lifetime between 0.1 and 10 ns, and significantly surpasses the reach of the LEP experiments. No significant excess above the background expectation is observed for candidate tracks with large transverse momentum, and constraints on chargino properties are obtained. In the AMSB scenarios, a chargino mass below 270 GeV is excluded at 95% confidence level.

DOI: [10.1103/PhysRevD.88.112006](https://doi.org/10.1103/PhysRevD.88.112006)

PACS numbers: 12.60.Jv, 13.85.Rm, 14.80.Nb

I. INTRODUCTION

Anomaly-mediated supersymmetry breaking (AMSB) models [1,2], where soft supersymmetry (SUSY) breaking is caused by loop effects, provide a constrained mass spectrum of SUSY particles. One prominent feature of these models is that the lightest supersymmetric particle is the nearly pure neutral wino that is mass degenerate with the charged wino. The lightest chargino ($\tilde{\chi}_1^\pm$) is then slightly heavier than the lightest neutralino ($\tilde{\chi}_1^0$) due to radiative corrections involving electroweak gauge bosons. The typical mass splitting between $\tilde{\chi}_1^\pm$ and $\tilde{\chi}_1^0$ ($\Delta m_{\tilde{\chi}_1}$) is ~ 160 MeV, which implies that $\tilde{\chi}_1^\pm$ has a considerable lifetime and predominantly decays into $\tilde{\chi}_1^0$ plus a low-momentum (~ 100 MeV) π^\pm . The mean lifetime ($\tau_{\tilde{\chi}_1^\pm}$) of $\tilde{\chi}_1^\pm$ is expressed in terms of $\Delta m_{\tilde{\chi}_1}$ and expected to be typically a fraction of a nanosecond. Several other SUSY models, which are motivated by the large value of the Higgs boson mass, also predict charginos with a significant lifetime and their decay to a soft pion and the lightest supersymmetric particle [3–6]. Therefore, some charginos could have decay lengths exceeding a few tens of centimeters at the Large Hadron Collider (LHC). When decaying in the sensitive volume, they are expected to be observed as “disappearing tracks” that have no more than a few associated hits in the outer region of the tracking system, and the softly emitted π^\pm is not reconstructed as it is curved away by the magnetic field. This article explores

AMSB scenarios by searching for charginos with their subsequent decays that result in such disappearing tracks. The electroweak production of charginos has a sizable cross section in proton-proton (pp) collisions at LHC energies. Chargino-pair and chargino-neutralino associated production processes are identified using jets of large transverse momentum (p_T) from initial-state radiation ($pp \rightarrow \tilde{\chi}_1^\pm \tilde{\chi}_1^0 j$ and $\tilde{\chi}_1^+ \tilde{\chi}_1^- j$, where j denotes a jet used to trigger the signal event). The search presented here, based on 20.3 fb^{-1} of 8 TeV pp collision data, increases the sensitivity compared to the previous ATLAS searches [7,8] due to analysis improvements and increases in the beam energy and luminosity. The most significant improvement is achieved by enhancing the track reconstruction efficiency for charginos having short decay lengths. In particular, the efficiency for charginos with $\tau_{\tilde{\chi}_1^\pm} \sim 0.2$ ns, predicted for $\Delta m_{\tilde{\chi}_1} \sim 160$ MeV, is around 100 times larger than in the previous searches. The present analysis also provides sensitivity to a wider range of chargino lifetimes and covers a larger angular acceptance. It significantly surpasses the reach of the LEP experiments [9–12] for charginos with lifetimes > 0.1 ns.

II. THE ATLAS DETECTOR

ATLAS is a multipurpose detector [13], covering nearly the entire solid angle [14] around the collision point with layers of tracking devices surrounded by a superconducting solenoid providing a 2 T axial magnetic field, a calorimeter system, and a muon spectrometer. The inner detector (ID) provides track reconstruction in the region $|\eta| < 2.5$ and consists of pixel and silicon microstrip (SCT) detectors inside a straw-tube transition radiation tracker (TRT). The pixel detector consists of three barrel layers and four disks in the forward and backward directions, providing on

*Full author list given at the end of the article.

Published by the American Physical Society under the terms of the [Creative Commons Attribution 3.0 License](https://creativecommons.org/licenses/by/3.0/). Further distribution of this work must maintain attribution to the author(s) and the published article's title, journal citation, and DOI.

average three measurement points for charged tracks. The SCT is composed of four cylindrical layers of double-sided silicon microstrip modules, with nine disk layers in each endcap region; eight silicon microstrip sensors are typically crossed by each track. The TRT, of particular importance to this search, covers $|\eta| < 1.0$ with its barrel detector, $0.8 < |\eta| < 2.0$ with the endcaps, and the radial range 563–1066 mm. The average number of TRT hits on a track going through the inner detector in the central region is about 32. Tracks in the transition region $0.8 < |\eta| < 1.2$ pass partially through both the barrel and endcap and are still expected to have >25 hits on average. Tracks passing through the dead region of the barrel TRT at $|\eta| < 0.1$ produce no TRT hits. The calorimeter system covers the range of $|\eta| < 4.9$. The electromagnetic calorimeter is a lead/liquid-argon (lead/LAr) detector in the barrel ($|\eta| < 1.475$) and endcap ($1.375 < |\eta| < 3.2$) regions. The hadronic calorimeters are composed of a steel and scintillator barrel ($|\eta| < 1.7$), a copper/LAr endcap ($1.5 < |\eta| < 3.2$), and a LAr forward system ($3.1 < |\eta| < 4.9$) with copper and tungsten absorbers. The muon spectrometer consists of three large superconducting toroids, trigger chambers, and precision tracking chambers that provide muon momentum measurements up to $|\eta| = 2.7$.

III. DATA AND SIMULATED EVENT SAMPLES

The data analyzed for this search were recorded in 2012 with the LHC colliding protons at $\sqrt{s} = 8$ TeV. The integrated luminosity, after the application of beam, detector, and data quality requirements, corresponds to $20.3 \pm 0.6 \text{ fb}^{-1}$, where the luminosity measurement is based on the calibration procedure described in Ref. [15] and uses the most recent van der Meer scans performed in November 2012 to determine the calibration and its uncertainty.

The analysis makes use of a dedicated topological trigger in order to suppress a huge Standard Model (SM) multijet background: it requires at least one jet with $p_T > 80$ GeV, large missing transverse momentum (its magnitude, E_T^{miss} , above 70 GeV), and $\Delta\phi_{\text{min}}^{\text{jet}-E_T^{\text{miss}}} > 1$, where $\Delta\phi_{\text{min}}^{\text{jet}-E_T^{\text{miss}}}$ indicates the azimuthal separation between the missing transverse momentum and the jet. If the event contains multiple jets with $p_T > 45$ GeV, the smallest $\Delta\phi_{\text{min}}^{\text{jet}-E_T^{\text{miss}}}$ value is taken by using either of the two highest- p_T jets. For the multijet background, $\Delta\phi_{\text{min}}^{\text{jet}-E_T^{\text{miss}}}$ peaks near zero since a large E_T^{miss} is usually due to jet mismeasurement and is thus aligned with a high- p_T jet, while the signal events cluster at $\Delta\phi_{\text{min}}^{\text{jet}-E_T^{\text{miss}}} \approx \pi$.

Simulated Monte Carlo (MC) events are used to assess the experimental sensitivity to given models. The minimal AMSB model is characterized by four parameters: the gravitino mass ($m_{3/2}$), the universal scalar mass (m_0), the ratio of Higgs vacuum expectation values at the

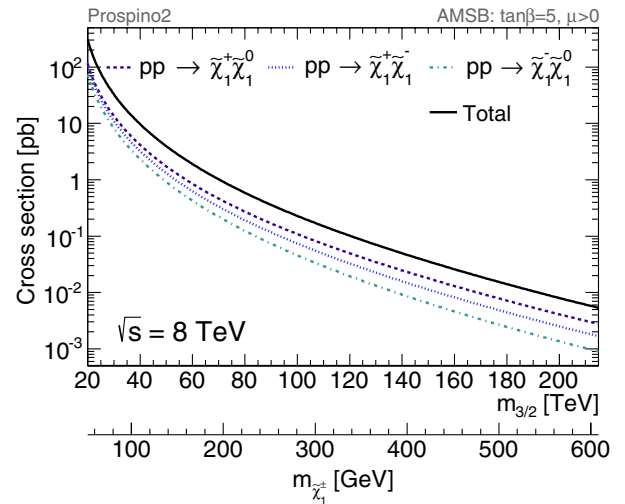


FIG. 1 (color online). The cross section for direct chargino production at $\sqrt{s} = 8$ TeV as a function of the gravitino mass $m_{3/2}$. The corresponding chargino mass $m_{\tilde{\chi}_1^\pm}$ for each $m_{3/2}$ value is indicated.

electroweak scale ($\tan\beta$), and the sign of the Higgsino mass term (μ). A large value of 1 TeV is used for m_0 in order to prevent the appearance of a tachyonic slepton. The production cross section is determined largely by the wino mass and is fairly independent of the other parameters. In this model, the wino mass is proportional to $m_{3/2}$. The SUSY mass spectrum and the decay tables are calculated with the ISASUSY from ISAJET v7.80 [16]. The corresponding MC signal samples are produced using Herwig++ 2.5.2 [17] with CTEQ6L1 [18] parton distribution functions of the proton (PDFs). All samples used in this article are produced using a detector simulation [19] based on GEANT4 [20] and include multiple pp interactions (pileup) in the triggered and adjacent bunch crossings to model the pileup effect. Simulated points with chargino masses ($m_{\tilde{\chi}_1^\pm}$) ranging from 80–600 GeV and various values of the chargino lifetime $\tau_{\tilde{\chi}_1^\pm}$ are generated. In the GEANT4 simulation the charginos decay exponentially and the branching fraction for the decay $\tilde{\chi}_1^\pm \rightarrow \tilde{\chi}_1^0 \pi^\pm$ is set to 100%. Signal cross sections are calculated at next-to-leading order in α_s using the PROSPINO2 [21] program as shown in Fig. 1. The nominal cross section and its uncertainty are taken from an envelope of cross section predictions using different PDF sets and factorization and renormalization scales, as described in Ref. [22].

IV. RECONSTRUCTION, OBJECT IDENTIFICATION, AND EVENT SELECTION

Standard Model processes, especially $W + \text{jet}$ events that naturally have large E_T^{miss} , can result in final-state kinematics similar to that of the signal. Kinematic selection criteria are applied to ensure high trigger efficiency and to reduce background arising from multijet processes

or from electroweak gauge bosons that decay leptonically. At the next stage, the vast majority of SM background events are removed by identifying and demanding a disappearing track in the event.

A. Track reconstruction

Charged particle trajectories are reconstructed as tracks in the ID. In order to improve the efficiency of reconstructing particles leaving short tracks, this analysis applies an extra extended-track reconstruction that provides pixel-seeded reconstructed tracks in addition to the ATLAS standard tracks. The standard track reconstruction algorithm [23] is a sequence made of two main steps. First, an inside-out sequence starts from triplets of three-dimensional space points from the pixel and SCT detectors, with each space point originating from a unique detector layer, and then extends the resulting trajectories by combining other pixel, SCT, and TRT hits. A second sequence takes the remaining TRT hits as seeds and attempts to extend identified trajectories inwards by combining them with unused space points. The first sequence is optimized to find primary tracks coming from the interaction point, while the second sequence is optimized for the reconstruction of electrons from photon conversions in the ID volume. The inside-out sequence is of particular interest for finding long-lived chargino trajectories, although it is optimized for the reconstruction of stable particles that leave long tracks in the ID, and, in particular, only reconstructs tracks with a minimum of seven space points. In order to increase the acceptance of the track reconstruction and especially the chargino track reconstruction efficiency at low radius, a third sequence is applied. This sequence proceeds using leftover pixel and SCT hits from the two previous tracking sequences and reconstructs tracks with a minimum of three pixel hits, while no SCT or TRT hits are required. The outward extension then follows; SCT and TRT hits are attached if they lie along the track trajectory. The tracks reconstructed by the third sequence are used only to select disappearing-track candidates.

B. Event reconstruction

The event vertex [24] is required to have at least five associated tracks. When more than one such vertex is found, the vertex with the largest $\sum |p_T|^2$ of the associated tracks is chosen as primary. Jets are reconstructed using the anti- k_r algorithm [25] with a distance parameter of 0.4. The inputs to the jet reconstruction algorithm are topological calorimeter energy clusters seeded by cells with energy significantly above the noise level. Jet energies are then calibrated back to the particle level [26]. Reconstructed jets must satisfy the requirements of $p_T > 20$ GeV and $|\eta| < 2.8$. Electron candidates are reconstructed from energy clusters in the electromagnetic calorimeter matched to a track in the ID. Electrons must then fulfill the “loose” identification requirements described in Ref. [27], have

transverse energy $E_T > 10$ GeV, and be within the region $|\eta| < 2.47$. Muon candidates are formed by matching ID tracks with either a complete track or a track segment reconstructed in the muon spectrometer [28]. Furthermore, muons must satisfy the requirements of $N_{\text{b-layer}} > 0$ if crossing an active module of the innermost pixel layer, $N_{\text{pixel}} > 0$, $N_{\text{SCT}} \geq 6$, $p_T > 10$ GeV, and $|\eta| < 2.4$, where $N_{\text{b-layer}}$, N_{pixel} , and N_{SCT} are the numbers of hits in the innermost pixel layer, the pixel and SCT detectors, respectively.

Following the object reconstruction described above, overlaps between jets and leptons are resolved to ensure isolation of leptons. First, any jet candidate lying within a distance of $\Delta R \equiv \sqrt{(\Delta\eta)^2 + (\Delta\phi)^2} = 0.2$ of an electron is discarded. Then, any lepton candidate within a distance of $\Delta R = 0.4$ of any surviving jet is discarded.

The calculation of E_T^{miss} is based on the transverse momenta of remaining jets and lepton candidates and on all calorimeter energy clusters that are not associated with such objects [29].

C. Kinematic selection

Following the event reconstruction, selection requirements to reject noncollision background events, given in Ref. [26], are applied to jets. In order to suppress backgrounds from $W/Z + \text{jets}$ and top-pair production processes, events are discarded if they contain any electron or muon candidates (lepton veto). Events containing muons are further suppressed by requiring no tracks with $p_T > 10$ GeV reconstructed in the muon spectrometer. The candidate events are finally required to have $E_T^{\text{miss}} > 90$ GeV, at least one jet with $p_T > 90$ GeV, and $\Delta\phi_{\text{min}}^{\text{jet-E}_T^{\text{miss}}} > 1.5$. The trigger selection is $>98\%$ efficient for signal events satisfying these selection requirements.

D. Selection of disappearing tracks

The tracks originating from charginos are expected to have high transverse momenta, to be isolated, and to have few associated hits in the outer region of the ID. The TRT detector, in particular, provides substantial discrimination against penetrating stable charged particles if only a small number of hits on the track is required. Therefore, candidate tracks for decaying charginos are required to fulfill the following criteria:

- (I) the track must have $N_{\text{pixel}} \geq 3$, $N_{\text{b-layer}} \geq 1$ if crossing an active module of the innermost pixel layer, $N_{\text{SCT}} \geq 2$, $|d_0| < 0.1$ mm, and $|z_0 \sin\theta| < 0.5$ mm, where d_0 and z_0 are the transverse and longitudinal impact parameters with respect to the primary vertex;
- (II) the track reconstruction must be of good quality, meeting the following requirements: it must have a track fit χ^2 probability of $>10\%$, no hits formed in a single pixel row of which the readout is shared with

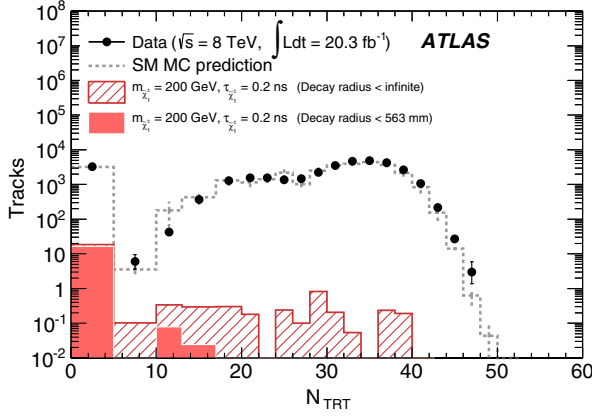


FIG. 2 (color online). Number of TRT hits (N_{TRT}) for data and signal MC events ($m_{\tilde{\chi}_1^\pm} = 200$ GeV, $\tau_{\tilde{\chi}_1^\pm} = 0.2$ ns) with the high- p_T isolated track selection. The expectation from SM MC events is also shown. The solid colored histogram shows the expected distribution for charginos with a decay radius < 563 mm while the hatched histogram shows it for all charginos for these mass and lifetime values. Tracks with $N_{\text{TRT}} < 5$ in SM events, mimicking the decaying-chargino signature, are described in Sec. V.

another pixel, and no hits missing in active silicon modules along the trajectory between the first and last hit of the track;

- (III) the track must be isolated: it must fulfill $p_T^{\text{cone40}}/p_T < 0.04$, where p_T^{cone40} is the sum of p_T of all tracks with $p_T > 400$ MeV, $|d_0| < 1.5$ mm, and $|z_0 \sin \theta| < 1.5$ mm that lie within a cone of $\Delta R = 0.4$ around the track. There must also be no jets having p_T above 45 GeV within a cone of $\Delta R = 0.4$ around the candidate track;
- (IV) the candidate track must have p_T above 15 GeV, and must be the highest- p_T isolated track in the event;
- (V) the candidate track must satisfy $0.1 < |\eta| < 1.9$;
- (VI) the number of TRT hits associated with the track (N_{TRT}), determined by counting hits lying on the extrapolated track, must be less than five.

Criteria (I) and (II) are applied in order to ensure well-reconstructed primary tracks. Criteria (III) and (IV) are employed to select chargino tracks that are isolated and have in most cases the highest p_T . These criteria also substantially reduce background tracks from the pileup. Criterion (V) is used to ensure coverage by the TRT active region and enhance the rejection of background tracks. Criterion (VI) helps to remove the majority of background tracks in SM processes, as shown in Fig. 2. For SM charged particles traversing the TRT detector, the number of TRT hits is typically $N_{\text{TRT}} \approx 32$, whereas for charginos that decay before reaching the TRT detector the expected is $N_{\text{TRT}} \approx 0$. Hereafter, “high- p_T isolated track selection” and “disappearing-track selection” indicate criteria (I)–(V) and (I)–(VI), respectively. Charginos triggered in this

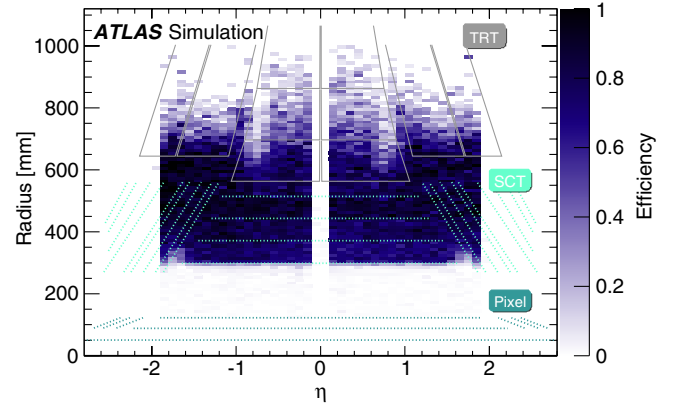


FIG. 3 (color online). The efficiency for decaying charginos with the disappearing-track selection. Vertical and horizontal axes are the radius and η of the decay, respectively. Sensitive layers and areas of the pixel, SCT, and TRT detectors are also indicated in the figure.

search should be highly boosted, thus no discriminant based on the energy loss dE/dx measurable in the pixel detector is adopted.

Making use of short tracks and the whole TRT detector for the background track rejection extends the sensitive decay volume inwards and enlarges the region of signal acceptance in η . This results in better sensitivity for charginos, especially with small lifetime, than in the previous search [8] based on the 7 TeV collision data. Figure 3 shows the tracking efficiency with the disappearing-track selection for decaying charginos as a function of the radius and η of the decay vertex. It is fully efficient for charginos that reach the first SCT layer and decay before reaching the TRT detector. The MC simulation shows that it is also largely independent of $m_{\tilde{\chi}_1^\pm}$. A summary of the kinematic selection criteria, disappearing-track requirements, and data reduction is given in Table I.

TABLE I. Summary of selection requirements and data reduction for data and expected signal events ($m_{\tilde{\chi}_1^\pm} = 200$ GeV, $\tau_{\tilde{\chi}_1^\pm} = 0.2$ ns). The signal selection efficiencies are also shown in parentheses. Signal efficiencies are low at the first stage due to the trigger based on a jet from initial-state radiation.

Selection requirement	Expected signal	
	Observed events	MC events (efficiency [%])
Quality requirements and trigger	20479553	1873 (8.8)
Jet cleaning	18627508	1867 (8.8)
Lepton veto	12485944	1827 (8.6)
Leading jet $p_T > 90$ GeV	10308840	1571 (7.4)
$E_T^{\text{miss}} > 90$ GeV	6113773	1484 (7.0)
$\Delta\phi_{\text{min}}^{\text{jet}-E_T^{\text{miss}}} > 1.5$	5604087	1444 (6.8)
High- p_T isolated track selection	34379	21.9 (0.10)
Disappearing-track selection	3256	18.4 (0.087)

V. ESTIMATE OF THE p_T SPECTRUM OF BACKGROUND TRACKS

There are three primary sources of tracks from background processes that mimic the disappearing-track signature: charged hadrons interacting with material in the ID (interacting-hadron tracks), prompt electrons or muons failing to satisfy their identification criteria (lepton tracks), and low- p_T charged particles whose p_T is highly mismeasured (p_T -mismeasured tracks). Interacting-hadron and electron tracks are responsible for the background in the approximate range $p_T < 50$ GeV, whereas p_T -mismeasured tracks are dominant for $p_T > 100$ GeV. A small contribution from muon tracks is expected throughout the full p_T range. The contribution of charged-hadron decays is significantly smaller than that of interacting hadrons; therefore, such a background source is neglected. A background estimation based on the MC simulation has difficulty accurately describing the properties of these background tracks. Therefore, the background contribution to the disappearing-track candidates is estimated using techniques that do not rely on the MC simulation. Each of the three types of background tracks shows a distinctive p_T spectrum; a simultaneous fit is performed for signal and background yields using the observed p_T spectrum and templates of background track p_T spectra produced from dedicated control data samples. The p_T spectra of the first two background types are obtained in the same way as in Ref. [8].

A. Interacting-hadron tracks

Charged hadrons, mostly charged pions, can interact with material in the ID and their tracks can be misidentified as disappearing tracks. The shape of the p_T distribution of interacting-hadron tracks is obtained from that of non-interacting-hadron tracks. In the p_T range above 15 GeV, where inelastic interactions dominate, the interaction rate has nearly no dependence on p_T [30], which is also confirmed by the detector simulation. By adopting kinematic selection criteria identical to those for the signal and ensuring traversal of the TRT detector by requiring $N_{\text{TRT}} > 25$, a data sample of non-interacting-hadron tracks is obtained. A pure control data sample is ensured by requiring associated calorimeter activity and removing the contamination from electron and muon tracks (described below) and any chargino signal. The following requirements are applied: $E_T^{\text{cone40}} > 7.5$ GeV and $\sum_{\Delta R < 0.4} E_T^{\text{clus}} / p_T^{\text{track}} > 0.4$, where E_T^{cone40} is the calorimeter transverse energy deposited in a cone of $\Delta R < 0.4$ around the track (excluding E_T of the calorimeter cluster matched to the track), $\sum_{\Delta R < 0.4} E_T^{\text{clus}}$ is the sum of cluster energies in a cone of $\Delta R < 0.4$ around the track, and p_T^{track} is the track p_T .

In most cases, interacting hadrons have associated calorimeter activity that can be used to form jets. Therefore, after the selection requirements, the contribution of this

background to the disappearing-track candidates having $p_T > 100$ GeV is negligibly small.

B. Leptons failing to satisfy identification criteria

Some charged leptons ($\ell \equiv e$ or μ) lose much of their momenta in the ID due to scattering with material or large bremsstrahlung. Such leptons are unlikely to be correctly identified (hence surviving the lepton veto) and may be classified as disappearing tracks.

In order to estimate the lepton-track background, a control data sample is defined by requiring kinematic selection identical to those for the signal search sample, while requiring one lepton that fulfills both its identification criteria and the isolated track selection criteria. The p_T spectrum of leptons without any identification requirements is obtained by applying a correction for the identification efficiency. The p_T distribution of lepton background tracks is then estimated by multiplying this distribution by the probability ($\mathcal{P}_\ell^{\text{dis}}$) of failing to satisfy the lepton identification criteria (hence being retained in the signal search sample) and passing the disappearing-track selection criteria. The electron and muon components are considered separately.

For the measurement of $\mathcal{P}_\ell^{\text{dis}}$, a tag-and-probe method is applied to $Z \rightarrow \ell\ell$ events collected with unprescaled single-lepton triggers and by requiring a Z boson candidate with reconstructed invariant mass within ± 5 GeV of the Z mass. Tag-leptons are required to be well isolated from jets and to fulfill the lepton identification criteria. Probe-leptons are selected without any identification requirements but with exactly the same high- p_T isolated track selection criteria used for chargino candidate tracks. The probability $\mathcal{P}_\ell^{\text{dis}}$ is given by the fraction of events in which the probe-lepton passes the disappearing-track selection criteria; it ranges between 10^{-2} and 10^{-4} for electrons and 10^{-4} and 10^{-5} for muons. Statistical uncertainties and uncertainties on the identification efficiency are considered in deriving the estimated p_T spectra and their uncertainties.

C. Tracks with mismeasured p_T

The background contribution to disappearing-track candidates with $p_T > 100$ GeV originates primarily from tracks with mismeasured p_T (p_T -mismeasured tracks). A high density of silicon hits, hadronic interactions, and scattering can lead to combinations of wrong space points in the procedure of track-seed finding or outward extension of trajectories, resulting in anomalously high values of p_T especially for short-length tracks. Simulation studies indicate that the p_T spectrum of such tracks depends little on the reconstructed d_0 or production process. Figure 4 shows the p_T spectrum of disappearing tracks with different d_0 values in a multijet-enriched data sample collected with single-jet triggers and requirements of $E_T^{\text{miss}} < 90$ GeV and no leptons: the contamination from

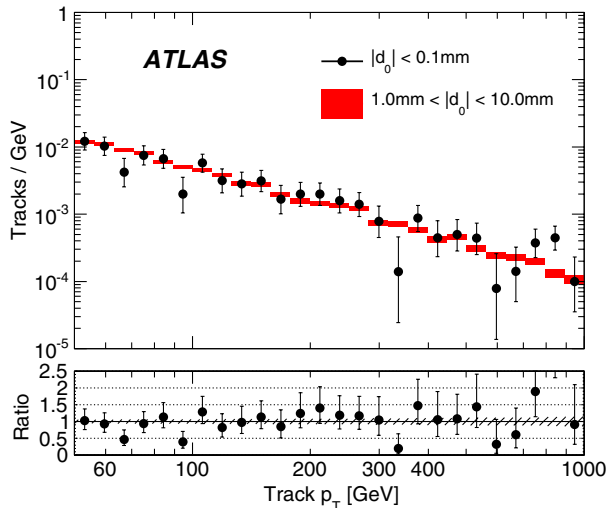


FIG. 4 (color online). The p_T distributions of disappearing tracks with impact parameter ranges $|d_0| < 0.1$ mm and $1 \text{ mm} < |d_0| < 10 \text{ mm}$ in the multijet-enriched data sample, normalized to unity. The ratio between the two distributions is also shown at the bottom of the figure. The error bars and band in the ratio plot indicate the statistical uncertainties of each sample.

interacting-hadron and lepton tracks is expected to be very small in the range $p_T > 50$ GeV or $|d_0| > 1$ mm. The p_T shape of p_T -mismeasured tracks with $|d_0| < 0.1$ mm is found to be the same as that of similarly mismeasured tracks with $1 \text{ mm} < |d_0| < 10 \text{ mm}$. A sample with a nearly pure p_T -mismeasured track contribution can be obtained with the same requirements as for the signal tracks, while requiring $1 \text{ mm} < |d_0| < 10 \text{ mm}$. The p_T shape is finally determined by a fit to the sample by a functional form x^{-a} ($x \equiv p_T^{\text{track}}$), where $a = 1.78 \pm 0.05$ is obtained.

VI. ESTIMATE OF SYSTEMATIC UNCERTAINTIES

The sources of systematic uncertainty on the signal expectation are the following: the theoretical cross section, parton radiation model, jet energy scale (JES) and resolution (JER), trigger efficiency, pileup modeling, track reconstruction efficiency, and the integrated luminosity. The contributions of each systematic uncertainty in the signal yield are summarized in Table II for two reference signal samples.

Theoretical uncertainties on the signal cross section, already described in Sec. III, range from 6% to 8% depending on $m_{\tilde{\chi}_1^\pm}$. The uncertainties on the modeling of high- p_T jets, originating from initial- and final-state radiation, are estimated by varying the generator tunes in the simulation as well as by generator-level studies carried out on samples produced with an additional jet in the matrix-element method using the MADGRAPH5 program [31] and the PYTHIA6 program [32]. By adopting PDF tunes that provide less and more radiation and taking the maximum deviation from the nominal tune, the uncertainty due to jet radiation

TABLE II. Summary of systematic uncertainties [%] on the expected number of signal events for $m_{\tilde{\chi}_1^\pm} = 200$ GeV and 300 GeV.

	200 GeV	300 GeV
(Theoretical uncertainty) Cross section	6.4	6.8
(Uncertainty on the acceptance)		
Modeling of initial/final-state radiation	14.5	16.4
JES/JER	3.9	6.0
Trigger efficiency	4.5	4.5
Pileup modeling	0.5	0.5
Track reconstruction efficiency	2.0	2.0
Luminosity	2.8	2.8
Subtotal	16.1	18.4

is evaluated. The uncertainty arising from the matching of matrix elements with parton showers is evaluated by doubling and halving the default value of the matching parameter [33]. The resulting changes are combined in quadrature and yield an uncertainty of 10%–17% depending on $m_{\tilde{\chi}_1^\pm}$. The uncertainties on the JES and JER result in a variation of the signal selection efficiency that is assessed according to Ref. [26], and an uncertainty of 3%–6% is assigned. An uncertainty due to the trigger efficiency is estimated to be 4.5% by taking the difference between data and MC simulation in a $W + \text{jet}$ sample in which W decays into μ plus ν_μ . The uncertainty originating from the pileup modeling in the simulation is evaluated by weighting simulated samples so that the average number of pileup interactions is varied by $\pm 10\%$, which yields a 0.5% uncertainty on the signal efficiency. The ID material affects the track reconstruction efficiency. An uncertainty of 2% is assigned from Ref. [34] to take into account differences in the tracking efficiency between data and MC simulation related to the detector material description in the simulation. The uncertainty on the integrated luminosity is $\pm 2.8\%$. It is derived, following the same methodology as that detailed in Ref. [15].

Systematic uncertainties on the background p_T shapes and normalizations arising from statistical uncertainties of the control data samples and uncertainties on the lepton identification efficiencies are also considered in deriving the results (discussed in Sec. VII). In order to account for a possible bias induced by the d_0 requirement in the control data sample of p_T -mismeasured tracks, an additional uncertainty is assigned by taking the difference between the value of the parameter a given in Sec. VC and the value 1.82 ± 0.07 derived using SM background MC events remaining after the selection requirements.

VII. FIT TO THE p_T SPECTRUM OF DISAPPEARING TRACKS

The signal hypothesis with a given value of $m_{\tilde{\chi}_1^\pm}$ and $\tau_{\tilde{\chi}_1^\pm}$ is tested based on an extended maximum likelihood fit to the p_T spectrum of the disappearing-track candidates. The

TABLE III. Numbers of observed and expected background events as well as the probability that a background-only experiment is more signal-like than observed (p_0) and the model-independent upper limit on the visible cross section ($\sigma_{\text{vis}}^{95\%}$) at 95% C.L.

	$p_T^{\text{track}} > 75 \text{ GeV}$	$p_T^{\text{track}} > 100 \text{ GeV}$	$p_T^{\text{track}} > 150 \text{ GeV}$	$p_T^{\text{track}} > 200 \text{ GeV}$
Observed events	59	36	19	13
Expected events	48.5 ± 12.3	37.1 ± 9.4	24.6 ± 6.3	18.0 ± 4.6
p_0 value	0.17	0.41	0.46	0.44
Observed $\sigma_{\text{vis}}^{95\%}$ [fb]	1.76	1.02	0.62	0.44
Expected $\sigma_{\text{vis}}^{95\%}$ [fb]	$1.42^{+0.50}_{-0.39}$	$1.05^{+0.37}_{-0.28}$	$0.67^{+0.27}_{-0.19}$	$0.56^{+0.23}_{-0.16}$

likelihood function for the track p_T consists of one probability density function for the signal and four for the different backgrounds derived in Sec. V. In the fit, the yields of the signal, interacting-hadron, and p_T -mismeasured tracks are left free. The yields of electron and muon background tracks are constrained to their estimated values within the uncertainties. The effects of systematic uncertainties on the yields and the parameters describing the p_T -distribution shapes of the background tracks are also incorporated into the likelihood function.

The number of observed events having a high- p_T disappearing track above a given threshold and the expectation for the background, derived by the background-only fit in the p_T range below 75 GeV, are given in Table III. No significant deviations from the background expectations are found. The probability (p_0 value) that a background-only experiment is more signal-like than the observation and the model-independent upper limit on the visible cross section ($\sigma_{\text{vis}}^{95\%}$) at 95% confidence level (C.L.) are also given in the table. Figure 5 shows the p_T distribution for

the selected data events compared to the background model derived by the background-only fit in the full p_T range: the best-fit values for the yields of interacting hadrons, electron tracks, muon tracks, and p_T -mismeasured tracks are 2187 ± 71 , 852 ± 35 , 23 ± 8 , and 212 ± 33 , respectively. Three selected examples for the signal are also shown in the figure.

An excess with a corresponding significance of $\sim 2\sigma$ is seen in Fig. 5 at p_T around 90 GeV. Detailed investigation of the events in this region show no peculiarities or significant differences in event kinematics or track properties compared to candidates in nearby track- p_T regions. The discrepancy is also not consistent with any of the signal hypotheses studied in this article. For the models considered, high- p_T tracks are expected and the best expected sensitivity derives from the region with p_T above 200 GeV, where a deficit is observed as reported in Table III.

Events with two disappearing-track candidates, being particularly sensitive to chargino-pair production with a long lifetime, are also explored. One candidate event is found; however, the event lacks high- p_T disappearing-track candidates (their p_T being 30 GeV and 18 GeV).

VIII. RESULTS

In the absence of a signal, constraints are set on $m_{\tilde{\chi}_1^\pm}$ and $\tau_{\tilde{\chi}_1^\pm}$. The upper limit on the production cross section for a given $m_{\tilde{\chi}_1^\pm}$ and $\tau_{\tilde{\chi}_1^\pm}$ at 95% C.L. is set at the point where the C.L. of the “signal + background” hypothesis, based on the profile likelihood ratio [35] and the CL_s prescription [36], falls below 5% when scanning the C.L. along various values of signal strength. The constraint on the allowed $\tau_{\tilde{\chi}_1^\pm} - m_{\tilde{\chi}_1^\pm}$ parameter space is shown in Fig. 6. The expected limit is set by the median of the distribution of 95% C.L. limits calculated by pseudoexperiments with the expected background and no signal, where the systematic parameters are varied according to their systematic uncertainties. The regions excluded by the previous ATLAS search [8] and the LEP2 searches are indicated. The example of the exclusion reached by the ALEPH experiment [9] of 8 GeV at 95% C.L. that is derived for the chargino mass in the case of heavy sfermions, irrespective of the chargino-neutralino mass difference, is shown as the

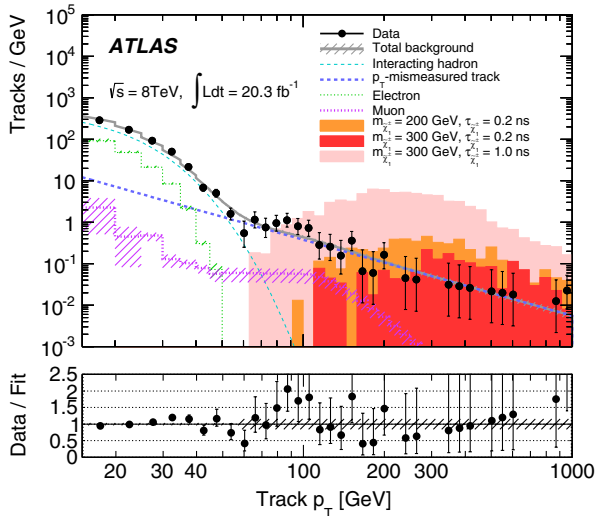


FIG. 5 (color online). The p_T distribution of disappearing-track candidates. The solid circles show data and lines show each background track p_T spectrum obtained by the background-only fit. The resulting uncertainties on the p_T spectrum for each background are indicated by the error bands. The signal expectations are also shown. The ratio of the data to the background track p_T spectrum is shown at the bottom of the figure.

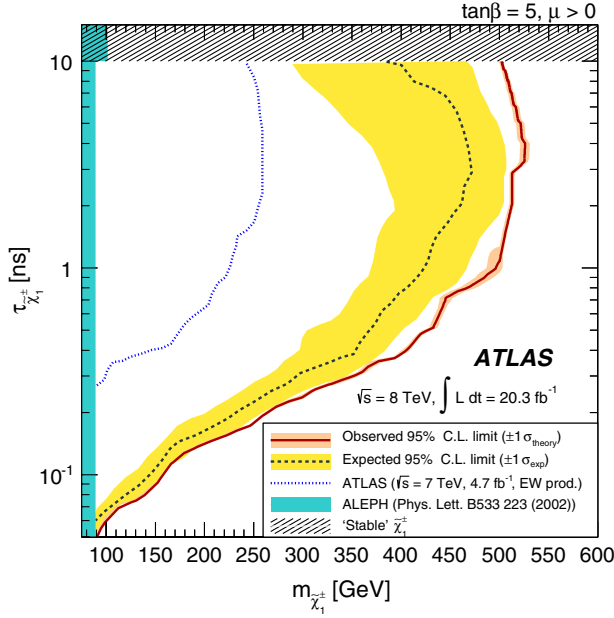


FIG. 6 (color online). The constraint on the allowed $\tau_{\tilde{\chi}_1^\pm} - m_{\tilde{\chi}_1^\pm}$ space for $\tan\beta = 5$ and $\mu > 0$. The black dashed line shows the expected limits at 95% C.L., with the surrounding shaded band indicating the 1σ exclusions due to experimental uncertainties. Observed limits are indicated by the solid bold contour representing the nominal limit and the narrow surrounding shaded band is obtained by varying the cross section by the theoretical scale and PDF uncertainties. The previous result from Ref. [8] and an example of the limits achieved at LEP2 by the ALEPH experiment [9] are also shown on the left by the dotted line and the shaded region, respectively. The search for charginos with long lifetimes, as indicated by the upper shaded region, is not covered by this analysis. The limits achieved at LEP2 by the ALEPH experiment of 101 GeV for long-lived charginos is taken from [9].

LEP2 result. This constraint is largely independent of $\tan\beta$ or the sign of μ .

The analysis is not performed for signals having $\tau_{\tilde{\chi}_1^\pm} > 10$ ns (corresponding $\Delta m_{\tilde{\chi}_1}$ being below the charged pion mass) because a significant fraction of charginos would traverse the ID before decaying, thereby reducing the event selection efficiency. In these scenarios the charginos are considered as stable particles and the main search tool would be to look for tracks with anomalous ionization energy loss [37]. In comparison with the previous result, the sensitivity to charginos having $\tau_{\tilde{\chi}_1^\pm} < 1$ ns is significantly improved and the exclusion reach is extended by ~ 200 GeV.

Figure 7 shows the constraint on the allowed $\Delta m_{\tilde{\chi}_1} - m_{\tilde{\chi}_1^\pm}$ parameter space of the minimal AMSB model; the expected 95% C.L. exclusion reaches $m_{\tilde{\chi}_1^\pm} = 245_{-30}^{+25}$ GeV for $\Delta m_{\tilde{\chi}_1} \sim 160$ MeV. The limits on $\tau_{\tilde{\chi}_1^\pm}$ are converted into limits on $\Delta m_{\tilde{\chi}_1}$ following Ref. [38]. The theoretical prediction of $\Delta m_{\tilde{\chi}_1}$ for winolike lightest chargino and

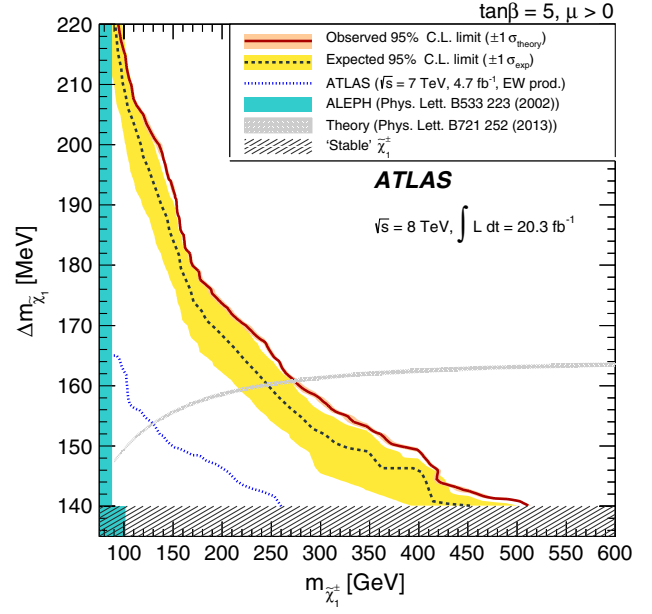


FIG. 7 (color online). The constraint on the allowed $\Delta m_{\tilde{\chi}_1} - m_{\tilde{\chi}_1^\pm}$ space of the AMSB model for $\tan\beta = 5$ and $\mu > 0$. The dashed line shows the expected limits at 95% C.L., with the surrounding shaded band indicating the 1σ exclusions due to experimental uncertainties. Observed limits are indicated by the solid bold contour representing the nominal limit and the narrow surrounding shaded band is obtained by varying the cross section by the theoretical scale and PDF uncertainties. The previous result from Ref. [8] and an example of the limits achieved at LEP2 by the ALEPH experiment [9] are also shown on the left by the dotted line and the shaded region, respectively. Charginos in the lower shaded region could have significantly longer lifetime values for which this analysis has no sensitivity as the chargino does not decay within the tracking volume. For this region of long-lived charginos, the limits achieved at LEP2 by the ALEPH experiment is 101 GeV [9].

neutralino states at two-loop level [39] is also indicated in the figure. A new limit that excludes charginos of $m_{\tilde{\chi}_1^\pm} < 270$ GeV (corresponding $\Delta m_{\tilde{\chi}_1}$ and $\tau_{\tilde{\chi}_1^\pm}$ being ~ 160 MeV and ~ 0.2 ns, respectively) at 95% C.L. is set in the AMSB models.

IX. CONCLUSIONS

The results from a search for charginos nearly mass degenerate with the lightest neutralino based on the high- p_T disappearing-track signature are presented. The analysis is based on 20.3 fb^{-1} of pp collisions at $\sqrt{s} = 8$ TeV collected by the ATLAS experiment at the LHC. The p_T spectrum of observed candidate tracks is found to be consistent with the expectation from SM background processes, and no indication of decaying charginos is observed. Constraints on the chargino mass, the mean lifetime, and the mass splitting are set, which are valid for most scenarios in which the lightest supersymmetric particle is a nearly pure neutral wino. In the AMSB

models, a chargino having a mass below 270 GeV is excluded at 95% C.L.

ACKNOWLEDGMENTS

We thank CERN for the very successful operation of the LHC, as well as the support staff from our institutions without whom ATLAS could not be operated efficiently. We acknowledge the support of ANPCyT, Argentina; YerPhI, Armenia; ARC, Australia; BMWF and FWF, Austria; ANAS, Azerbaijan; SSTC, Belarus; CNPq and FAPESP, Brazil; NSERC, NRC and CFI, Canada; CERN; CONICYT, Chile; CAS, MOST, and NSFC, China; COLCIENCIAS, Colombia; MSMT CR, MPO CR, and VSC CR, Czech Republic; DNRF, DNSRC, and Lundbeck Foundation, Denmark; EPLANET, ERC, and NSRF, European Union; IN2P3-CNRS, CEA-DSM/IRFU, France; GNSF, Georgia; BMBF, DFG, HGF, MPG, and AvH Foundation, Germany; GSRT and NSRF, Greece; ISF,

MINERVA, GIF, DIP, and Benoziyo Center, Israel; INFN, Italy; MEXT and JSPS, Japan; CNRST, Morocco; FOM and NWO, Netherlands; BRF and RCN, Norway; MNiSW, Poland; GRICES and FCT, Portugal; MNE/IFA, Romania; MES of Russia and ROSATOM, Russian Federation; JINR; MSTD, Serbia; MSSR, Slovakia; ARRS and MIZŠ, Slovenia; DST/NRF, South Africa; MINECO, Spain; SRC and Wallenberg Foundation, Sweden; SER, SNSF, and Cantons of Bern and Geneva, Switzerland; NSC, Taiwan; TAEK, Turkey; STFC, the Royal Society and Leverhulme Trust, United Kingdom; DOE and NSF, United States of America. The crucial computing support from all WLCG partners is acknowledged gratefully, in particular, from CERN and the ATLAS Tier-1 facilities at TRIUMF (Canada), NDGF (Denmark, Norway, Sweden), CC-IN2P3 (France), KIT/GridKA (Germany), INFN-CNAF (Italy), NL-T1 (Netherlands), PIC (Spain), ASGC (Taiwan), RAL (UK) and BNL (USA), and in the Tier-2 facilities worldwide.

-
- [1] G. F. Giudice, M. A. Luty, H. Murayama, and R. Rattazzi, *J. High Energy Phys.* **12** (1998) 027.
- [2] L. Randall and R. Sundrum, *Nucl. Phys.* **B557**, 79 (1999).
- [3] M. Ibe, S. Matsumoto, and T. T. Yanagida, *Phys. Rev. D* **85**, 095011 (2012).
- [4] L. J. Hall, Y. Nomura, and S. Shirai, *J. High Energy Phys.* **01** (2013) 036.
- [5] A. Arvanitaki, N. Craig, S. Dimopoulos, and G. Villadoro, *J. High Energy Phys.* **02** (2013) 126.
- [6] N. Arkani-Hamed, A. Gupta, D. E. Kaplan, N. Weiner, and T. Zorawski, [arXiv:1212.6971](https://arxiv.org/abs/1212.6971).
- [7] ATLAS Collaboration, *Eur. Phys. J. C* **72**, 1993 (2012).
- [8] ATLAS Collaboration, *J. High Energy Phys.* **01** (2013) 131.
- [9] A. Heister *et al.* (ALEPH Collaboration), *Phys. Lett. B* **533**, 223 (2002).
- [10] J. Abdallah *et al.* (DELPHI Collaboration), *Eur. Phys. J. C* **34**, 145 (2004).
- [11] M. Acciarri *et al.* (L3 Collaboration), *Phys. Lett. B* **482**, 31 (2000).
- [12] G. Abbiendi *et al.* (OPAL Collaboration), *Eur. Phys. J. C* **26**, 479 (2003).
- [13] ATLAS Collaboration, *JINST* **3**, S08003 (2008).
- [14] ATLAS uses a right-handed coordinate system with its origin at the nominal interaction point (IP) in the center of the detector and the z axis coinciding with the axis of the beam pipe. The x axis points from the IP to the center of the LHC ring, and the y axis points upward. Cylindrical coordinates (r, ϕ) are used in the transverse plane, ϕ being the azimuthal angle around the beam pipe. Pseudorapidity is defined in terms of the polar angle θ as $\eta = -\ln \tan(\theta/2)$.
- [15] ATLAS Collaboration, *Eur. Phys. J. C* **73**, 2518 (2013).
- [16] F. E. Paige, S. D. Protopopescu, H. Baer, and X. Tata, [arXiv:hep-ph/0312045](https://arxiv.org/abs/hep-ph/0312045).
- [17] M. Bahr *et al.*, *Eur. Phys. J. C* **58**, 639 (2008).
- [18] J. Pumplin, D. R. Stump, J. Huston, H.-L. Lai, P. Nadolsky, and W.-K. Tung, *J. High Energy Phys.* **07** (2002) 012.
- [19] ATLAS Collaboration, *Eur. Phys. J. C* **70**, 823 (2010).
- [20] S. Agostinelli *et al.* (GEANT4 Collaboration), *Nucl. Instrum. Methods Phys. Res., Sect. A* **506**, 250 (2003).
- [21] W. Beenakker, R. Hopker, M. Spira, and P. Zerwas, *Nucl. Phys.* **B492**, 51 (1997).
- [22] M. Kramer *et al.*, [arXiv:1206.2892](https://arxiv.org/abs/1206.2892).
- [23] T. Cornelisen, M. Elsing, I. Gavrilenko, W. Liebig, E. Moyses, and A. Salzburger, *J. Phys. Conf. Ser.* **119**, 032014 (2008).
- [24] ATLAS Collaboration, Report No. ATLAS-CONF-2010-069, <http://cdsweb.cern.ch/record/1281344>.
- [25] M. Cacciari, G. P. Salam, and G. Soyez, *J. High Energy Phys.* **04** (2008) 063.
- [26] ATLAS Collaboration, *Eur. Phys. J. C* **73**, 2304 (2013).
- [27] ATLAS Collaboration, *Eur. Phys. J. C* **72**, 1909 (2012).
- [28] ATLAS Collaboration, Report No. ATLAS-CONF-2011-063, <https://cds.cern.ch/record/1345743>.
- [29] ATLAS Collaboration, *Eur. Phys. J. C* **72**, 1844 (2012).
- [30] K. Nakamura *et al.* (Particle Data Group Collaboration), *J. Phys. G* **37**, 075021 (2010). Corresponding data may be found at pp. 398–399.
- [31] J. Alwall, M. Herquet, F. Maltoni, O. Mattelaer, and T. Stelzer, *J. High Energy Phys.* **06** (2011) 128.
- [32] T. Sjöstrand, S. Mrenna, and P. Z. Skands, *J. High Energy Phys.* **05** (2006) 026.

- [33] J. Alwall *et al.*, *Eur. Phys. J. C* **53**, 473 (2008).
 [34] ATLAS Collaboration, *Eur. Phys. J. C* **71**, 1636 (2011).
 [35] G. Cowan, K. Cranmer, E. Gross, and O. Vitells, *Eur. Phys. J. C* **71**, 1554 (2011).
 [36] A.L. Read, *J. Phys. G* **28**, 2693 (2002).
 [37] ATLAS Collaboration, *Phys. Lett. B* **720**, 277 (2013).
 [38] C. Chen, M. Drees, and J. Gunion, [arXiv:hep-ph/9902309](https://arxiv.org/abs/hep-ph/9902309).
 [39] M. Ibe, S. Matsumoto, and R. Sato, *Phys. Lett. B* **721**, 252 (2013).

G. Aad,⁴⁸ T. Abajyan,²¹ B. Abbott,¹¹² J. Abdallah,¹² S. Abdel Khalek,¹¹⁶ O. Abdinov,¹¹ R. Aben,¹⁰⁶ B. Abi,¹¹³ M. Abolins,⁸⁹ O. S. AbouZeid,¹⁵⁹ H. Abramowicz,¹⁵⁴ H. Abreu,¹³⁷ Y. Abulaiti,^{147a,147b} B. S. Acharya,^{165a,165b,b} L. Adamczyk,^{38a} D. L. Adams,²⁵ T. N. Addy,⁵⁶ J. Adelman,¹⁷⁷ S. Adomeit,⁹⁹ T. Adye,¹³⁰ S. Aefsky,²³ T. Agatonovic-Jovin,^{13b} J. A. Aguilar-Saavedra,^{125b,c} M. Agustoni,¹⁷ S. P. Ahlen,²² A. Ahmad,¹⁴⁹ F. Ahmadov,^{64,d} M. Ahsan,⁴¹ G. Aielli,^{134a,134b} T. P. A. Åkesson,⁸⁰ G. Akimoto,¹⁵⁶ A. V. Akimov,⁹⁵ M. A. Alam,⁷⁶ J. Albert,¹⁷⁰ S. Albrand,⁵⁵ M. J. Alconada Verzini,⁷⁰ M. Aleksa,³⁰ I. N. Aleksandrov,⁶⁴ F. Alessandria,^{90a} C. Alexa,^{26a} G. Alexander,¹⁵⁴ G. Alexandre,⁴⁹ T. Alexopoulos,¹⁰ M. Alhroob,^{165a,165c} M. Aliev,¹⁶ G. Alimonti,^{90a} L. Alio,⁸⁴ J. Alison,³¹ B. M. M. Allbrooke,¹⁸ L. J. Allison,⁷¹ P. P. Allport,⁷³ S. E. Allwood-Spiers,⁵³ J. Almond,⁸³ A. Aloisio,^{103a,103b} R. Alon,¹⁷³ A. Alonso,³⁶ F. Alonso,⁷⁰ A. Altheimer,³⁵ B. Alvarez Gonzalez,⁸⁹ M. G. Alviggi,^{103a,103b} K. Amako,⁶⁵ Y. Amaral Coutinho,^{24a} C. Amelung,²³ V. V. Ammosov,^{129,a} S. P. Amor Dos Santos,^{125a} A. Amorim,^{125a,e} S. Amoroso,⁴⁸ N. Amram,¹⁵⁴ G. Amundsen,²³ C. Anastopoulos,³⁰ L. S. Ancu,¹⁷ N. Andari,³⁰ T. Andeen,³⁵ C. F. Anders,^{58b} G. Anders,^{58a} K. J. Anderson,³¹ A. Andreazza,^{90a,90b} V. Andrei,^{58a} X. S. Anduaga,⁷⁰ S. Angelidakis,⁹ P. Anger,⁴⁴ A. Angerami,³⁵ F. Anghinolfi,³⁰ A. V. Anisenkov,¹⁰⁸ N. Anjos,^{125a} A. Annovi,⁴⁷ A. Antonaki,⁹ M. Antonelli,⁴⁷ A. Antonov,⁹⁷ J. Antos,^{145b} F. Anulli,^{133a} M. Aoki,¹⁰² L. Aperio Bella,¹⁸ R. Apolle,^{119,f} G. Arabidze,⁸⁹ I. Aracena,¹⁴⁴ Y. Arai,⁶⁵ A. T. H. Arce,⁴⁵ S. Arfaoui,¹⁴⁹ J-F. Arguin,⁹⁴ S. Argyropoulos,⁴² E. Arik,^{19a,a} M. Arik,^{19a} A. J. Armbruster,⁸⁸ O. Arnaez,⁸² V. Arnal,⁸¹ O. Arslan,²¹ A. Artamonov,⁹⁶ G. Artoni,^{133a,133b} S. Asai,¹⁵⁶ N. Asbah,⁹⁴ S. Ask,²⁸ B. Åsman,^{147a,147b} L. Asquith,⁶ K. Assamagan,²⁵ R. Astalos,^{145a} A. Astbury,¹⁷⁰ M. Atkinson,¹⁶⁶ N. B. Atlay,¹⁴² B. Auerbach,⁶ E. Auge,¹¹⁶ K. Augsten,¹²⁷ M. Auresseau,^{146b} G. Avolio,³⁰ G. Azuelos,^{94,g} Y. Azuma,¹⁵⁶ M. A. Baak,³⁰ C. Bacci,^{135a,135b} A. M. Bach,¹⁵ H. Bachacou,¹³⁷ K. Bachas,¹⁵⁵ M. Backes,³⁰ M. Backhaus,²¹ J. Backus Mayes,¹⁴⁴ E. Badescu,^{26a} P. Bagiacchi,^{133a,133b} P. Bagnaia,^{133a,133b} Y. Bai,^{33a} D. C. Bailey,¹⁵⁹ T. Bain,³⁵ J. T. Baines,¹³⁰ O. K. Baker,¹⁷⁷ S. Baker,⁷⁷ P. Balek,¹²⁸ F. Balli,¹³⁷ E. Banas,³⁹ Sw. Banerjee,¹⁷⁴ D. Banfi,³⁰ A. Bangert,¹⁵¹ V. Bansal,¹⁷⁰ H. S. Bansil,¹⁸ L. Barak,¹⁷³ S. P. Baranov,⁹⁵ T. Barber,⁴⁸ E. L. Barberio,⁸⁷ D. Barberis,^{50a,50b} M. Barbero,⁸⁴ D. Y. Bardin,⁶⁴ T. Barillari,¹⁰⁰ M. Barisonzi,¹⁷⁶ T. Barklow,¹⁴⁴ N. Barlow,²⁸ B. M. Barnett,¹³⁰ R. M. Barnett,¹⁵ A. Baroncelli,^{135a} G. Barone,⁴⁹ A. J. Barr,¹¹⁹ F. Barreiro,⁸¹ J. Barreiro Guimarães da Costa,⁵⁷ R. Bartoldus,¹⁴⁴ A. E. Barton,⁷¹ V. Bartsch,¹⁵⁰ A. Bassalat,¹¹⁶ A. Basye,¹⁶⁶ R. L. Bates,⁵³ L. Batkova,^{145a} J. R. Batley,²⁸ M. Battistin,³⁰ F. Bauer,¹³⁷ H. S. Bawa,^{144,h} T. Beau,⁷⁹ P. H. Beauchemin,¹⁶² R. Beccherle,^{50a} P. Bechtel,²¹ H. P. Beck,¹⁷ K. Becker,¹⁷⁶ S. Becker,⁹⁹ M. Beckingham,¹³⁹ A. J. Beddall,^{19c} A. Beddall,^{19c} S. Bedikian,¹⁷⁷ V. A. Bednyakov,⁶⁴ C. P. Bee,⁸⁴ L. J. Beamster,¹⁰⁶ T. A. Beermann,¹⁷⁶ M. Begel,²⁵ K. Behr,¹¹⁹ C. Belanger-Champagne,⁸⁶ P. J. Bell,⁴⁹ W. H. Bell,⁴⁹ G. Bella,¹⁵⁴ L. Bellagamba,^{20a} A. Bellerive,²⁹ M. Bellomo,³⁰ A. Belloni,⁵⁷ O. L. Beloborodova,^{108,i} K. Belotskiy,⁹⁷ O. Beltramello,³⁰ O. Benary,¹⁵⁴ D. Bencheikroun,^{136a} K. Bendtz,^{147a,147b} N. Benekos,¹⁶⁶ Y. Benhammou,¹⁵⁴ E. Benhar Noccioli,⁴⁹ J. A. Benitez Garcia,^{160b} D. P. Benjamin,⁴⁵ J. R. Bensinger,²³ K. Benslama,¹³¹ S. Bentvelsen,¹⁰⁶ D. Berge,³⁰ E. Bergeas Kuutmann,¹⁶ N. Berger,⁵ F. Berghaus,¹⁷⁰ E. Berglund,¹⁰⁶ J. Beringer,¹⁵ C. Bernard,²² P. Bernat,⁷⁷ R. Bernhard,⁴⁸ C. Bernius,⁷⁸ F. U. Bernlochner,¹⁷⁰ T. Berry,⁷⁶ P. Berta,¹²⁸ C. Bertella,⁸⁴ F. Bertolucci,^{123a,123b} M. I. Besana,^{90a} G. J. Besjes,¹⁰⁵ O. Bessidskaia,^{147a,147b} N. Besson,¹³⁷ S. Bethke,¹⁰⁰ W. Bhimji,⁴⁶ R. M. Bianchi,¹²⁴ L. Bianchini,²³ M. Bianco,³⁰ O. Biebel,⁹⁹ S. P. Bieniek,⁷⁷ K. Bierwagen,⁵⁴ J. Biesiada,¹⁵ M. Biglietti,^{135a} J. Bilbao De Mendizabal,⁴⁹ H. Bilokon,⁴⁷ M. Bindi,^{20a,20b} S. Binet,¹¹⁶ A. Bingul,^{19c} C. Bini,^{133a,133b} B. Bittner,¹⁰⁰ C. W. Black,¹⁵¹ J. E. Black,¹⁴⁴ K. M. Black,²² D. Blackburn,¹³⁹ R. E. Blair,⁶ J.-B. Blanchard,¹³⁷ T. Blazek,^{145a} I. Bloch,⁴² C. Blocker,²³ J. Blocki,³⁹ W. Blum,^{82,a} U. Blumenschein,⁵⁴ G. J. Bobbink,¹⁰⁶ V. S. Bobrovnikov,¹⁰⁸ S. S. Bocchetta,⁸⁰ A. Bocci,⁴⁵ C. R. Boddy,¹¹⁹ M. Boehler,⁴⁸ J. Boek,¹⁷⁶ T. T. Boek,¹⁷⁶ N. Boelaert,³⁶ J. A. Bogaerts,³⁰ A. G. Bogdanchikov,¹⁰⁸ A. Bogouch,^{91,a} C. Bohm,^{147a} J. Bohm,¹²⁶ V. Boisvert,⁷⁶ T. Bold,^{38a} V. Boldea,^{26a} A. S. Boldyrev,⁹⁸ N. M. Bolnet,¹³⁷ M. Bomben,⁷⁹ M. Bona,⁷⁵ M. Boonekamp,¹³⁷ S. Bordononi,⁷⁹ C. Borer,¹⁷ A. Borisov,¹²⁹ G. Borissov,⁷¹ M. Borri,⁸³ S. Borroni,⁴² J. Bortfeldt,⁹⁹ V. Bortolotto,^{135a,135b} K. Bos,¹⁰⁶ D. Boscherini,^{20a} M. Bosman,¹² H. Boterenbrood,¹⁰⁶ J. Bouchami,⁹⁴ J. Boudreau,¹²⁴ E. V. Bouhova-Thacker,⁷¹ D. Boumediene,³⁴ C. Bourdarios,¹¹⁶ N. Bousson,⁸⁴ S. Boutouil,^{136d}

- A. Boveia,³¹ J. Boyd,³⁰ I. R. Boyko,⁶⁴ I. Bozovic-Jelisavcic,^{13b} J. Bracinik,¹⁸ P. Branchini,^{135a} A. Brandt,⁸ G. Brandt,¹⁵ O. Brandt,⁵⁴ U. Bratzler,¹⁵⁷ B. Brau,⁸⁵ J. E. Brau,¹¹⁵ H. M. Braun,^{176,a} S. F. Brazzale,^{165a,165c} B. Brelief,¹⁵⁹ K. Brendlinger,¹²¹ R. Brenner,¹⁶⁷ S. Bressler,¹⁷³ T. M. Bristow,⁴⁶ D. Britton,⁵³ F. M. Brochu,²⁸ I. Brock,²¹ R. Brock,⁸⁹ F. Broggi,^{90a} C. Bromberg,⁸⁹ J. Bronner,¹⁰⁰ G. Brooijmans,³⁵ T. Brooks,⁷⁶ W. K. Brooks,^{32b} J. Brosamer,¹⁵ E. Brost,¹¹⁵ G. Brown,⁸³ J. Brown,⁵⁵ P. A. Bruckman de Renstrom,³⁹ D. Bruncko,^{145b} R. Bruneliere,⁴⁸ S. Brunet,⁶⁰ A. Bruni,^{20a} G. Bruni,^{20a} M. Bruschi,^{20a} L. Bryngemark,⁸⁰ T. Buanes,¹⁴ Q. Buat,⁵⁵ F. Bucci,⁴⁹ J. Buchanan,¹¹⁹ P. Buchholz,¹⁴² R. M. Buckingham,¹¹⁹ A. G. Buckley,⁴⁶ S. I. Buda,^{26a} I. A. Budagov,⁶⁴ B. Budick,¹⁰⁹ F. Buehrer,⁴⁸ L. Bugge,¹¹⁸ O. Bulekov,⁹⁷ A. C. Bundock,⁷³ M. Bunse,⁴³ H. Burckhart,³⁰ S. Burdin,⁷³ T. Burgess,¹⁴ S. Burke,¹³⁰ I. Burmeister,⁴³ E. Busato,³⁴ V. Büscher,⁸² P. Bussey,⁵³ C. P. Buszello,¹⁶⁷ B. Butler,⁵⁷ J. M. Butler,²² A. I. Butt,³ C. M. Buttar,⁵³ J. M. Butterworth,⁷⁷ W. Buttinger,²⁸ A. Buzatu,⁵³ M. Byszewski,¹⁰ S. Cabrera Urbán,¹⁶⁸ D. Caforio,^{20a,20b} O. Cakir,^{4a} P. Calafiura,¹⁵ G. Calderini,⁷⁹ P. Calfayan,⁹⁹ R. Calkins,¹⁰⁷ L. P. Caloba,^{24a} R. Caloi,^{133a,133b} D. Calvet,³⁴ S. Calvet,³⁴ R. Camacho Toro,⁴⁹ P. Camarri,^{134a,134b} D. Cameron,¹¹⁸ L. M. Caminada,¹⁵ R. Caminal Armadans,¹² S. Campana,³⁰ M. Campanelli,⁷⁷ V. Canale,^{103a,103b} F. Canelli,³¹ A. Canepa,^{160a} J. Cantero,⁸¹ R. Cantrill,⁷⁶ T. Cao,⁴⁰ M. D. M. Capeans Garrido,³⁰ I. Caprini,^{26a} M. Caprini,^{26a} M. Capua,^{37a,37b} R. Caputo,⁸² R. Cardarelli,^{134a} T. Carli,³⁰ G. Carlino,^{103a} L. Carminati,^{90a,90b} S. Caron,¹⁰⁵ E. Carquin,^{32a} G. D. Carrillo-Montoya,^{146c} A. A. Carter,⁷⁵ J. R. Carter,²⁸ J. Carvalho,^{125a,j} D. Casadei,⁷⁷ M. P. Casado,¹² C. Caso,^{50a,50b,a} E. Castaneda-Miranda,^{146b} A. Castelli,¹⁰⁶ V. Castillo Gimenez,¹⁶⁸ N. F. Castro,^{125a} P. Catastini,⁵⁷ A. Catinaccio,³⁰ J. R. Catmore,⁷¹ A. Cattai,³⁰ G. Cattani,^{134a,134b} S. Caughron,⁸⁹ V. Cavaliere,¹⁶⁶ D. Cavalli,^{90a} M. Cavalli-Sforza,¹² V. Cavasinni,^{123a,123b} F. Ceradini,^{135a,135b} B. Cerio,⁴⁵ A. S. Cerqueira,^{24b} A. Cerri,¹⁵ L. Cerrito,⁷⁵ F. Cerutti,¹⁵ A. Cervelli,¹⁷ S. A. Cetin,^{19b} A. Chafaq,^{136a} D. Chakraborty,¹⁰⁷ I. Chalupkova,¹²⁸ K. Chan,³ P. Chang,¹⁶⁶ B. Chapleau,⁸⁶ J. D. Chapman,²⁸ J. W. Chapman,⁸⁸ D. Charfeddine,¹¹⁶ D. G. Charlton,¹⁸ V. Chavda,⁸³ C. A. Chavez Barajas,³⁰ S. Cheatham,⁸⁶ S. Chekanov,⁶ S. V. Chekulaev,^{160a} G. A. Chelkov,⁶⁴ M. A. Chelstowska,⁸⁸ C. Chen,⁶³ H. Chen,²⁵ K. Chen,¹⁴⁹ S. Chen,^{33c} X. Chen,¹⁷⁴ Y. Chen,³⁵ Y. Cheng,³¹ A. Cheplakov,⁶⁴ R. Cherkaoui El Moursli,^{136e} V. Chernyatin,^{25,a} E. Cheu,⁷ L. Chevalier,¹³⁷ V. Chiarella,⁴⁷ G. Chiefari,^{103a,103b} J. T. Childers,³⁰ A. Chilingarov,⁷¹ G. Chiodini,^{72a} A. S. Chisholm,¹⁸ R. T. Chislett,⁷⁷ A. Chitan,^{26a} M. V. Chizhov,⁶⁴ G. Choudalakis,³¹ S. Chouridou,⁹ B. K. B. Chow,⁹⁹ I. A. Christidi,⁷⁷ D. Chromek-Burckhart,³⁰ M. L. Chu,¹⁵² J. Chudoba,¹²⁶ G. Ciapetti,^{133a,133b} A. K. Ciftci,^{4a} R. Ciftci,^{4a} D. Cinca,⁶² V. Cindro,⁷⁴ A. Ciocio,¹⁵ M. Cirilli,⁸⁸ P. Cirkovic,^{13b} Z. H. Citron,¹⁷³ M. Citterio,^{90a} M. Ciubancan,^{26a} A. Clark,⁴⁹ P. J. Clark,⁴⁶ R. N. Clarke,¹⁵ J. C. Clemens,⁸⁴ B. Clement,⁵⁵ C. Clement,^{147a,147b} Y. Coadou,⁸⁴ M. Cobal,^{165a,165c} A. Coccaro,¹³⁹ J. Cochran,⁶³ S. Coelli,^{90a} L. Coffey,²³ J. G. Cogan,¹⁴⁴ J. Coggeshall,¹⁶⁶ J. Colas,⁵ B. Cole,³⁵ S. Cole,¹⁰⁷ A. P. Colijn,¹⁰⁶ C. Collins-Tooth,⁵³ J. Collot,⁵⁵ T. Colombo,^{58c} G. Colon,⁸⁵ G. Compostella,¹⁰⁰ P. Conde Muiño,^{125a} E. Coniavitis,¹⁶⁷ M. C. Conidi,¹² S. M. Consonni,^{90a,90b} V. Consorti,⁴⁸ S. Constantinescu,^{26a} C. Conta,^{120a,120b} G. Conti,⁵⁷ F. Conventi,^{103a,k} M. Cooke,¹⁵ B. D. Cooper,⁷⁷ A. M. Cooper-Sarkar,¹¹⁹ N. J. Cooper-Smith,⁷⁶ K. Copic,¹⁵ T. Cornelissen,¹⁷⁶ M. Corradi,^{20a} F. Corriveau,^{86,l} A. Corso-Radu,¹⁶⁴ A. Cortes-Gonzalez,¹² G. Cortiana,¹⁰⁰ G. Costa,^{90a} M. J. Costa,¹⁶⁸ D. Costanzo,¹⁴⁰ D. Côté,⁸ G. Cottin,^{32a} L. Courneyea,¹⁷⁰ G. Cowan,⁷⁶ B. E. Cox,⁸³ K. Cranmer,¹⁰⁹ G. Cree,²⁹ S. Crépe-Renaudin,⁵⁵ F. Crescioli,⁷⁹ M. Cristinziani,²¹ G. Crosetti,^{37a,37b} C.-M. Cuciuc,^{26a} C. Cuenca Almenar,¹⁷⁷ T. Cuhadar Donszelmann,¹⁴⁰ J. Cummings,¹⁷⁷ M. Curatolo,⁴⁷ C. Cuthbert,¹⁵¹ H. Czirr,¹⁴² P. Czodrowski,⁴⁴ Z. Czyczula,¹⁷⁷ S. D'Auria,⁵³ M. D'Onofrio,⁷³ A. D'Orazio,^{133a,133b} M. J. Da Cunha Sargedas De Sousa,^{125a} C. Da Via,⁸³ W. Dabrowski,^{38a} A. Dafinca,¹¹⁹ T. Dai,⁸⁸ F. Dallaire,⁹⁴ C. Dallapiccola,⁸⁵ M. Dam,³⁶ D. S. Damiani,¹³⁸ A. C. Daniells,¹⁸ M. Dano Hoffmann,³⁶ V. Dao,¹⁰⁵ G. Darbo,^{50a} G. L. Darlea,^{26c} S. Darmora,⁸ J. A. Dassoulas,⁴² W. Davey,²¹ C. David,¹⁷⁰ T. Davidek,¹²⁸ E. Davies,^{119,f} M. Davies,⁹⁴ O. Davignon,⁷⁹ A. R. Davison,⁷⁷ Y. Davygora,^{58a} E. Dawe,¹⁴³ I. Dawson,¹⁴⁰ R. K. Daya-Ishmukhametova,²³ K. De,⁸ R. de Asmundis,^{103a} S. De Castro,^{20a,20b} S. De Cecco,⁷⁹ J. de Graat,⁹⁹ N. De Groot,¹⁰⁵ P. de Jong,¹⁰⁶ C. De La Taille,¹¹⁶ H. De la Torre,⁸¹ F. De Lorenzi,⁶³ L. De Nooij,¹⁰⁶ D. De Pedis,^{133a} A. De Salvo,^{133a} U. De Sanctis,^{165a,165c} A. De Santo,¹⁵⁰ J. B. De Vivie De Regie,¹¹⁶ G. De Zorzi,^{133a,133b} W. J. Dearnaley,⁷¹ R. Debbe,²⁵ C. Debenedetti,⁴⁶ B. Dechenaux,⁵⁵ D. V. Dedovich,⁶⁴ J. Degenhardt,¹²¹ J. Del Peso,⁸¹ T. Del Prete,^{123a,123b} T. Delemontex,⁵⁵ F. Deliot,¹³⁷ M. Deliyergiyev,⁷⁴ A. Dell'Acqua,³⁰ L. Dell'Asta,²² M. Della Pietra,^{103a,k} D. della Volpe,^{103a,103b} M. Delmastro,⁵ P. A. Delsart,⁵⁵ C. Deluca,¹⁰⁶ S. Demers,¹⁷⁷ M. Demichev,⁶⁴ A. Demilly,⁷⁹ B. Demirkoz,^{12,m} S. P. Denisov,¹²⁹ D. Derendarz,³⁹ J. E. Derkaoui,^{136d} F. Derue,⁷⁹ P. Dervan,⁷³ K. Desch,²¹ P. O. Deviveiros,¹⁰⁶ A. Dewhurst,¹³⁰ B. DeWilde,¹⁴⁹ S. Dhaliwal,¹⁰⁶ R. Dhullipudi,^{78,n} A. Di Ciaccio,^{134a,134b} L. Di Ciaccio,⁵ C. Di Donato,^{103a,103b} A. Di Girolamo,³⁰

B. Di Girolamo,³⁰ A. Di Mattia,¹⁵³ B. Di Micco,^{135a,135b} R. Di Nardo,⁴⁷ A. Di Simone,⁴⁸ R. Di Sipio,^{20a,20b}
D. Di Valentino,²⁹ M. A. Diaz,^{32a} E. B. Diehl,⁸⁸ J. Dietrich,⁴² T. A. Dietzsch,^{58a} S. Diglio,⁸⁷ K. Dindar Yagci,⁴⁰
J. Dingfelder,²¹ C. Dionisi,^{133a,133b} P. Dita,^{26a} S. Dita,^{26a} F. Dittus,³⁰ F. Djama,⁸⁴ T. Djobava,^{51b} M. A. B. do Vale,^{24c}
A. Do Valle Wemans,^{125a,o} T. K. O. Doan,⁵ D. Dobos,³⁰ E. Dobson,⁷⁷ J. Dodd,³⁵ C. Doglioni,⁴⁹ T. Doherty,⁵³
T. Dohmae,¹⁵⁶ Y. Doi,^{65,a} J. Dolejsi,¹²⁸ Z. Dolezal,¹²⁸ B. A. Dolgoshein,^{97,a} M. Donadelli,^{24d} S. Donati,^{123a,123b}
J. Donini,³⁴ J. Dopke,³⁰ A. Doria,^{103a} A. Dos Anjos,¹⁷⁴ A. Dotti,^{123a,123b} M. T. Dova,⁷⁰ A. T. Doyle,⁵³ M. Dris,¹⁰
J. Dubbert,⁸⁸ S. Dube,¹⁵ E. Dubreuil,³⁴ E. Duchovni,¹⁷³ G. Duckeck,⁹⁹ O. A. Ducu,^{26a} D. Duda,¹⁷⁶ A. Dudarev,³⁰
F. Dudziak,⁶³ L. Dufflot,¹¹⁶ L. Duguid,⁷⁶ M. Dührssen,³⁰ M. Dunford,^{58a} H. Duran Yildiz,^{4a} M. Düren,⁵²
M. Dwuznik,^{38a} J. Ebke,⁹⁹ W. Edson,² C. A. Edwards,⁷⁶ N. C. Edwards,⁴⁶ W. Ehrenfeld,²¹ T. Eifert,¹⁴⁴ G. Eigen,¹⁴
K. Einsweiler,¹⁵ E. Eisenhandler,⁷⁵ T. Ekelof,¹⁶⁷ M. El Kacimi,^{136c} M. Ellert,¹⁶⁷ S. Elles,⁵ F. Ellinghaus,⁸² K. Ellis,⁷⁵
N. Ellis,³⁰ J. Elmsheuser,⁹⁹ M. Elsing,³⁰ D. Emeliyanov,¹³⁰ Y. Enari,¹⁵⁶ O. C. Endner,⁸² M. Endo,¹¹⁷
R. Engelmann,¹⁴⁹ J. Erdmann,¹⁷⁷ A. Ereditato,¹⁷ D. Eriksson,^{147a} G. Ernis,¹⁷⁶ J. Ernst,² M. Ernst,²⁵ J. Ernwein,¹³⁷
D. Errede,¹⁶⁶ S. Errede,¹⁶⁶ E. Ertel,⁸² M. Escalier,¹¹⁶ H. Esch,⁴³ C. Escobar,¹²⁴ X. Espinal Curull,¹² B. Esposito,⁴⁷
F. Etienne,⁸⁴ A. I. Etienne,¹³⁷ E. Etzion,¹⁵⁴ D. Evangelakou,⁵⁴ H. Evans,⁶⁰ L. Fabbri,^{20a,20b} G. Facini,³⁰
R. M. Fakhruddinov,¹²⁹ S. Falciano,^{133a} Y. Fang,^{33a} M. Fanti,^{90a,90b} A. Farbin,⁸ A. Farilla,^{135a} T. Farooque,¹⁵⁹
S. Farrell,¹⁶⁴ S. M. Farrington,¹⁷¹ P. Farthouat,³⁰ F. Fassi,¹⁶⁸ P. Fassnacht,³⁰ D. Fassouliotis,⁹ B. Fatholahzadeh,¹⁵⁹
A. Favareto,^{50a,50b} L. Fayard,¹¹⁶ P. Federic,^{145a} O. L. Fedin,¹²² W. Fedorko,¹⁶⁹ M. Fehling-Kaschek,⁴⁸ L. Felgioni,⁸⁴
C. Feng,^{33d} E. J. Feng,⁶ H. Feng,⁸⁸ A. B. Fenyuk,¹²⁹ W. Fernando,⁶ S. Ferrag,⁵³ J. Ferrando,⁵³ V. Ferrara,⁴²
A. Ferrari,¹⁶⁷ P. Ferrari,¹⁰⁶ R. Ferrari,^{120a} D. E. Ferreira de Lima,⁵³ A. Ferrer,¹⁶⁸ D. Ferrere,⁴⁹ C. Ferretti,⁸⁸
A. Ferretto Parodi,^{50a,50b} M. Fiascaris,³¹ F. Fiedler,⁸² A. Filipčič,⁷⁴ M. Filipuzzi,⁴² F. Filthaut,¹⁰⁵
M. Fincke-Keeler,¹⁷⁰ K. D. Finelli,⁴⁵ M. C. N. Fiolhais,^{125a,j} L. Fiorini,¹⁶⁸ A. Firan,⁴⁰ J. Fischer,¹⁷⁶ M. J. Fisher,¹¹⁰
E. A. Fitzgerald,²³ M. Flechl,⁴⁸ I. Fleck,¹⁴² P. Fleischmann,¹⁷⁵ S. Fleischmann,¹⁷⁶ G. T. Fletcher,¹⁴⁰ G. Fletcher,⁷⁵
T. Flick,¹⁷⁶ A. Floderus,⁸⁰ L. R. Flores Castillo,¹⁷⁴ A. C. Florez Bustos,^{160b} M. J. Flowerdew,¹⁰⁰ T. Fonseca Martin,¹⁷
A. Formica,¹³⁷ A. Forti,⁸³ D. Fortin,^{160a} D. Fournier,¹¹⁶ H. Fox,⁷¹ P. Francavilla,¹² M. Franchini,^{20a,20b}
S. Franchino,³⁰ D. Francis,³⁰ M. Franklin,⁵⁷ S. Franz,⁶¹ M. Fraternali,^{120a,120b} S. Fratina,¹²¹ S. T. French,²⁸
C. Friedrich,⁴² F. Friedrich,⁴⁴ D. Froidevaux,³⁰ J. A. Frost,²⁸ C. Fukunaga,¹⁵⁷ E. Fullana Torregrosa,¹²⁸
B. G. Fulson,¹⁴⁴ J. Fuster,¹⁶⁸ C. Gabaldon,⁵⁵ O. Gabizon,¹⁷³ A. Gabrielli,^{20a,20b} A. Gabrielli,^{133a,133b} S. Gadatsch,¹⁰⁶
T. Gadfort,²⁵ S. Gadomski,⁴⁹ G. Gagliardi,^{50a,50b} P. Gagnon,⁶⁰ C. Galea,⁹⁹ B. Galhardo,^{125a} E. J. Gallas,¹¹⁹ V. Gallo,¹⁷
B. J. Gallop,¹³⁰ P. Gallus,¹²⁷ G. Galster,³⁶ K. K. Gan,¹¹⁰ R. P. Gandrajula,⁶² J. Gao,^{33b,p} Y. S. Gao,^{144,h}
F. M. Garay Walls,⁴⁶ F. Garbersson,¹⁷⁷ C. García,¹⁶⁸ J. E. García Navarro,¹⁶⁸ M. Garcia-Sciveres,¹⁵ R. W. Gardner,³¹
N. Garelli,¹⁴⁴ V. Garonne,³⁰ C. Gatti,⁴⁷ G. Gaudio,^{120a} B. Gaur,¹⁴² L. Gauthier,⁹⁴ P. Gauzzi,^{133a,133b}
I. L. Gavrilenko,⁹⁵ C. Gay,¹⁶⁹ G. Gaycken,²¹ E. N. Gazis,¹⁰ P. Ge,^{33d,q} Z. Gecse,¹⁶⁹ C. N. P. Gee,¹³⁰
D. A. A. Geerts,¹⁰⁶ Ch. Geich-Gimbel,²¹ K. Gellerstedt,^{147a,147b} C. Gemme,^{50a} A. Gemmell,⁵³ M. H. Genest,⁵⁵
S. Gentile,^{133a,133b} M. George,⁵⁴ S. George,⁷⁶ D. Gerbaudo,¹⁶⁴ A. Gershon,¹⁵⁴ H. Ghazlane,^{136b} N. Ghodbane,³⁴
B. Giacobbe,^{20a} S. Giagu,^{133a,133b} V. Giangiobbe,¹² P. Giannetti,^{123a,123b} F. Gianotti,³⁰ B. Gibbard,²⁵ S. M. Gibson,⁷⁶
M. Gilchriese,¹⁵ T. P. S. Gillam,²⁸ D. Gillberg,³⁰ A. R. Gillman,¹³⁰ D. M. Gingrich,^{3,g} N. Giokaris,⁹
M. P. Giordani,^{165a,165c} R. Giordano,^{103a,103b} F. M. Giorgi,¹⁶ P. Giovannini,¹⁰⁰ P. F. Giraud,¹³⁷ D. Giugni,^{90a}
C. Giuliani,⁴⁸ M. Giunta,⁹⁴ B. K. Gjelsten,¹¹⁸ I. Gkialas,^{155,r} L. K. Gladilin,⁹⁸ C. Glasman,⁸¹ J. Glatzer,²¹
A. Glazov,⁴² G. L. Glonti,⁶⁴ M. Goblirsch-Kolb,¹⁰⁰ J. R. Goddard,⁷⁵ J. Godfrey,¹⁴³ J. Godlewski,³⁰ C. Goeringer,⁸²
S. Goldfarb,⁸⁸ T. Golling,¹⁷⁷ D. Golubkov,¹²⁹ A. Gomes,^{125a,e} L. S. Gomez Fajardo,⁴² R. Gonçalves,⁷⁶
J. Goncalves Pinto Firmino Da Costa,⁴² L. Gonella,²¹ S. González de la Hoz,¹⁶⁸ G. Gonzalez Parra,¹²
M. L. Gonzalez Silva,²⁷ S. Gonzalez-Sevilla,⁴⁹ J. J. Goodson,¹⁴⁹ L. Goossens,³⁰ P. A. Gorbounov,⁹⁶ H. A. Gordon,²⁵
I. Gorelov,¹⁰⁴ G. Gorfine,¹⁷⁶ B. Gorini,³⁰ E. Gorini,^{72a,72b} A. Gorišek,⁷⁴ E. Gornicki,³⁹ A. T. Goshaw,⁶ C. Gössling,⁴³
M. I. Gostkin,⁶⁴ I. Gough Eschrich,¹⁶⁴ M. Gouighri,^{136a} D. Goujdami,^{136c} M. P. Goulette,⁴⁹ A. G. Goussiou,¹³⁹
C. Goy,⁵ S. Gozpinar,²³ H. M. X. Grabas,¹³⁷ L. Graber,⁵⁴ I. Grabowska-Bold,^{38a} P. Grafström,^{20a,20b} K.-J. Grahm,⁴²
J. Gramling,⁴⁹ E. Gramstad,¹¹⁸ F. Grancagnolo,^{72a} S. Grancagnolo,¹⁶ V. Grassi,¹⁴⁹ V. Gratchev,¹²² H. M. Gray,³⁰
J. A. Gray,¹⁴⁹ E. Graziani,^{135a} O. G. Grebenyuk,¹²² Z. D. Greenwood,^{78,n} K. Gregersen,³⁶ I. M. Gregor,⁴²
P. Grenier,¹⁴⁴ J. Griffiths,⁸ N. Grigalashvili,⁶⁴ A. A. Grillo,¹³⁸ K. Grimm,⁷¹ S. Grinstein,^{12,s} Ph. Gris,³⁴
Y. V. Grishkevich,⁹⁸ J.-F. Grivaz,¹¹⁶ J. P. Grohs,⁴⁴ A. Grohsjean,⁴² E. Gross,¹⁷³ J. Grosse-Knetter,⁵⁴
G. C. Grossi,^{134a,134b} J. Groth-Jensen,¹⁷³ Z. J. Grout,¹⁵⁰ K. Grybel,¹⁴² F. Guescini,⁴⁹ D. Guest,¹⁷⁷ O. Gueta,¹⁵⁴
C. Guicheney,³⁴ E. Guido,^{50a,50b} T. Guillemin,¹¹⁶ S. Guindon,² U. Gul,⁵³ C. Gumpert,⁴⁴ J. Gunther,¹²⁷ J. Guo,³⁵

- S. Gupta,¹¹⁹ P. Gutierrez,¹¹² N. G. Gutierrez Ortiz,⁵³ C. Gutsche,⁷⁷ N. Guttman,¹⁵⁴ C. Guyot,¹³⁷ C. Gwenlan,¹¹⁹
 C. B. Gwilliam,⁷³ A. Haas,¹⁰⁹ C. Haber,¹⁵ H. K. Hadavand,⁸ P. Haefner,²¹ S. Hageboeck,²¹ Z. Hajduk,³⁹
 H. Hakobyan,¹⁷⁸ D. Hall,¹¹⁹ G. Halladjian,⁶² K. Hamacher,¹⁷⁶ P. Hamal,¹¹⁴ K. Hamano,⁸⁷ M. Hamer,⁵⁴
 A. Hamilton,^{146a,t} S. Hamilton,¹⁶² L. Han,^{33b} K. Hanagaki,¹¹⁷ K. Hanawa,¹⁵⁶ M. Hance,¹⁵ C. Handel,⁸² P. Hanke,^{58a}
 J. R. Hansen,³⁶ J. B. Hansen,³⁶ J. D. Hansen,³⁶ P. H. Hansen,³⁶ P. Hansson,¹⁴⁴ K. Hara,¹⁶¹ A. S. Hard,¹⁷⁴
 T. Harenberg,¹⁷⁶ S. Harkusha,⁹¹ D. Harper,⁸⁸ R. D. Harrington,⁴⁶ O. M. Harris,¹³⁹ P. F. Harrison,¹⁷¹ F. Hartjes,¹⁰⁶
 A. Harvey,⁵⁶ S. Hasegawa,¹⁰² Y. Hasegawa,¹⁴¹ S. Hassani,¹³⁷ S. Haug,¹⁷ M. Hauschild,³⁰ R. Hauser,⁸⁹
 M. Havranek,²¹ C. M. Hawkes,¹⁸ R. J. Hawkins,³⁰ A. D. Hawkins,⁸⁰ T. Hayashi,¹⁶¹ D. Hayden,⁸⁹ C. P. Hays,¹¹⁹
 H. S. Hayward,⁷³ S. J. Haywood,¹³⁰ S. J. Head,¹⁸ T. Heck,⁸² V. Hedberg,⁸⁰ L. Heelan,⁸ S. Heim,¹²¹ B. Heinemann,¹⁵
 S. Heisterkamp,³⁶ J. Hejbal,¹²⁶ L. Helary,²² C. Heller,⁹⁹ M. Heller,³⁰ S. Hellman,^{147a,147b} D. Hellmich,²¹
 C. Hensels,³⁰ J. Henderson,¹¹⁹ R. C. W. Henderson,⁷¹ A. Henrichs,¹⁷⁷ A. M. Henriques Correia,³⁰
 S. Henrot-Versille,¹¹⁶ C. Hensel,⁵⁴ G. H. Herbert,¹⁶ C. M. Hernandez,⁸ Y. Hernández Jiménez,¹⁶⁸
 R. Herrberg-Schubert,¹⁶ G. Herten,⁴⁸ R. Hertenberger,⁹⁹ L. Hervas,³⁰ G. G. Hesketh,⁷⁷ N. P. Hessey,¹⁰⁶ R. Hickling,⁷⁵
 E. Higón-Rodríguez,¹⁶⁸ J. C. Hill,²⁸ K. H. Hiller,⁴² S. Hillert,²¹ S. J. Hillier,¹⁸ I. Hinchliffe,¹⁵ E. Hines,¹²¹
 M. Hirose,¹¹⁷ D. Hirschbuehl,¹⁷⁶ J. Hobbs,¹⁴⁹ N. Hod,¹⁰⁶ M. C. Hodgkinson,¹⁴⁰ P. Hodgson,¹⁴⁰ A. Hoecker,³⁰
 M. R. Hoferkamp,¹⁰⁴ J. Hoffman,⁴⁰ D. Hoffmann,⁸⁴ J. I. Hofmann,^{58a} M. Hohlfeld,⁸² T. R. Holmes,¹⁵
 S. O. Holmgren,^{147a} T. M. Hong,¹²¹ L. Hooft van Huysduynen,¹⁰⁹ J.-Y. Hostachy,⁵⁵ S. Hou,¹⁵² A. Hoummada,^{136a}
 J. Howard,¹¹⁹ J. Howarth,⁸³ M. Hrabovsky,¹¹⁴ I. Hristova,¹⁶ J. Hrivnac,¹¹⁶ T. Hryn'ova,⁵ P. J. Hsu,⁸² S.-C. Hsu,¹³⁹
 D. Hu,³⁵ X. Hu,²⁵ Y. Huang,^{146c} Z. Hubacek,³⁰ F. Hubaut,⁸⁴ F. Huegging,²¹ A. Huettmann,⁴² T. B. Huffman,¹¹⁹
 E. W. Hughes,³⁵ G. Hughes,⁷¹ M. Huhtinen,³⁰ T. A. Hülsing,⁸² M. Hurwitz,¹⁵ N. Huseynov,^{64,d} J. Huston,⁸⁹ J. Huth,⁵⁷
 G. Iacobucci,⁴⁹ G. Iakovidis,¹⁰ I. Ibragimov,¹⁴² L. Iconomidou-Fayard,¹¹⁶ J. Idarraga,¹¹⁶ E. Ideal,¹⁷⁷ P. Iengo,^{103a}
 O. Igonkina,¹⁰⁶ T. Iizawa,¹⁷² Y. Ikegami,⁶⁵ K. Ikematsu,¹⁴² M. Ikeno,⁶⁵ D. Iliadis,¹⁵⁵ N. Ilic,¹⁵⁹ Y. Inamaru,⁶⁶
 T. Ince,¹⁰⁰ P. Ioannou,⁹ M. Iodice,^{135a} K. Iordanidou,⁹ V. Ippolito,^{133a,133b} A. Irls Quiles,¹⁶⁸ C. Isaksson,¹⁶⁷
 M. Ishino,⁶⁷ M. Ishitsuka,¹⁵⁸ R. Ishmukhametov,¹¹⁰ C. Issever,¹¹⁹ S. Istin,^{19a} A. V. Ivashin,¹²⁹ W. Iwanski,³⁹
 H. Iwasaki,⁶⁵ J. M. Izen,⁴¹ V. Izzo,^{103a} B. Jackson,¹²¹ J. N. Jackson,⁷³ M. Jackson,⁷³ P. Jackson,¹ M. R. Jaekel,³⁰
 V. Jain,² K. Jakobs,⁴⁸ S. Jakobsen,³⁶ T. Jakoubek,¹²⁶ J. Jakubek,¹²⁷ D. O. Jamin,¹⁵² D. K. Jana,¹¹² E. Jansen,⁷⁷
 H. Jansen,³⁰ J. Janssen,²¹ M. Janus,¹⁷¹ R. C. Jared,¹⁷⁴ G. Jarlskog,⁸⁰ L. Jeanty,⁵⁷ G.-Y. Jeng,¹⁵¹ I. Jen-La Plante,³¹
 D. Jennens,⁸⁷ P. Jenni,^{48,u} J. Jentzsch,⁴³ C. Jeske,¹⁷¹ S. Jézéquel,⁵ M. K. Jha,^{20a} H. Ji,¹⁷⁴ W. Ji,⁸² J. Jia,¹⁴⁹ Y. Jiang,^{33b}
 M. Jimenez Belenguer,⁴² S. Jin,^{33a} A. Jinaru,^{26a} O. Jinnouchi,¹⁵⁸ M. D. Joergensen,³⁶ D. Joffe,⁴⁰ K. E. Johansson,^{147a}
 P. Johansson,¹⁴⁰ K. A. Johns,⁷ K. Jon-And,^{147a,147b} G. Jones,¹⁷¹ R. W. L. Jones,⁷¹ T. J. Jones,⁷³ P. M. Jorge,^{125a}
 K. D. Joshi,⁸³ J. Jovicevic,¹⁴⁸ X. Ju,¹⁷⁴ C. A. Jung,⁴³ R. M. Jungst,³⁰ P. Jussel,⁶¹ A. Juste Rozas,^{12,s} M. Kaci,¹⁶⁸
 A. Kaczmarska,³⁹ P. Kadlecik,³⁶ M. Kado,¹¹⁶ H. Kagan,¹¹⁰ M. Kagan,¹⁴⁴ E. Kajomovitz,⁴⁵ S. Kalinin,¹⁷⁶ S. Kama,⁴⁰
 N. Kanaya,¹⁵⁶ M. Kaneda,³⁰ S. Kaneti,²⁸ T. Kanno,¹⁵⁸ V. A. Kantserov,⁹⁷ J. Kanzaki,⁶⁵ B. Kaplan,¹⁰⁹ A. Kapliy,³¹
 D. Kar,⁵³ K. Karakostas,¹⁰ N. Karastathis,¹⁰ M. Karnevskiy,⁸² S. N. Karpov,⁶⁴ K. Karthik,¹⁰⁹ V. Kartvelishvili,⁷¹
 A. N. Karyukhin,¹²⁹ L. Kashif,¹⁷⁴ G. Kasieczka,^{58b} R. D. Kass,¹¹⁰ A. Kastanas,¹⁴ Y. Kataoka,¹⁵⁶ A. Katre,⁴⁹
 J. Katzy,⁴² V. Kaushik,⁷ K. Kawagoe,⁶⁹ T. Kawamoto,¹⁵⁶ G. Kawamura,⁵⁴ S. Kazama,¹⁵⁶ V. F. Kazanin,¹⁰⁸
 M. Y. Kazarinov,⁶⁴ R. Keeler,¹⁷⁰ P. T. Keener,¹²¹ R. Kehoe,⁴⁰ M. Keil,⁵⁴ J. S. Keller,¹³⁹ H. Keoshkerian,⁵
 O. Kepka,¹²⁶ B. P. Kerševan,⁷⁴ S. Kersten,¹⁷⁶ K. Kessoku,¹⁵⁶ J. Keung,¹⁵⁹ F. Khalil-zada,¹¹ H. Khandanyan,^{147a,147b}
 A. Khanov,¹¹³ D. Kharchenko,⁶⁴ A. Khodinov,⁹⁷ A. Khomich,^{58a} T. J. Khoo,²⁸ G. Khoriali,²¹ A. Khoroshilov,¹⁷⁶
 V. Khovanskiy,⁹⁶ E. Khramov,⁶⁴ J. Khubua,^{51b} H. Kim,^{147a,147b} S. H. Kim,¹⁶¹ N. Kimura,¹⁷² O. Kind,¹⁶ B. T. King,⁷³
 M. King,⁶⁶ R. S. B. King,¹¹⁹ S. B. King,¹⁶⁹ J. Kirk,¹³⁰ A. E. Kiryunin,¹⁰⁰ T. Kishimoto,⁶⁶ D. Kisielevska,^{38a}
 T. Kitamura,⁶⁶ T. Kittelmann,¹²⁴ K. Kiuchi,¹⁶¹ E. Kladiva,^{145b} M. Klein,⁷³ U. Klein,⁷³ K. Kleinknecht,⁸²
 P. Klimek,^{147a,147b} A. Klimentov,²⁵ R. Klingenberg,⁴³ J. A. Klinger,⁸³ E. B. Klinkby,³⁶ T. Klioutchnikova,³⁰
 P. F. Klok,¹⁰⁵ E.-E. Kluge,^{58a} P. Kluit,¹⁰⁶ S. Kluth,¹⁰⁰ E. Kneringer,⁶¹ E. B. F. G. Knoops,⁸⁴ A. Knue,⁵⁴ B. R. Ko,⁴⁵
 T. Kobayashi,¹⁵⁶ M. Kobel,⁴⁴ M. Kocian,¹⁴⁴ P. Kodys,¹²⁸ S. Koenig,⁸² P. Koevesarki,²¹ T. Koffas,²⁹ E. Koffeman,¹⁰⁶
 L. A. Kogan,¹¹⁹ S. Kohlmann,¹⁷⁶ Z. Kohout,¹²⁷ T. Kohriki,⁶⁵ T. Koi,¹⁴⁴ H. Kolanoski,¹⁶ I. Koletsou,⁵ J. Koll,⁸⁹
 A. A. Komar,^{95,a} Y. Komori,¹⁵⁶ T. Kondo,⁶⁵ K. Köneke,⁴⁸ A. C. König,¹⁰⁵ T. Kono,^{65,v} R. Konoplich,^{109,w}
 N. Konstantinidis,⁷⁷ R. Kopeliansky,¹⁵³ S. Koperny,^{38a} L. Köpke,⁸² A. K. Kopp,⁴⁸ K. Korcyl,³⁹ K. Kordas,¹⁵⁵
 A. Korn,⁴⁶ A. A. Korol,¹⁰⁸ I. Korolkov,¹² E. V. Korolkova,¹⁴⁰ V. A. Korotkov,¹²⁹ O. Kortner,¹⁰⁰ S. Kortner,¹⁰⁰
 V. V. Kostyukhin,²¹ S. Kotov,¹⁰⁰ V. M. Kotov,⁶⁴ A. Kotwal,⁴⁵ C. Kourkoumelis,⁹ V. Kouskoura,¹⁵⁵ A. Koutsman,^{160a}
 R. Kowalewski,¹⁷⁰ T. Z. Kowalski,^{38a} W. Kozanecki,¹³⁷ A. S. Kozhin,¹²⁹ V. Kral,¹²⁷ V. A. Kramarenko,⁹⁸

G. Kramberger,⁷⁴ M. W. Krasny,⁷⁹ A. Krasznahorkay,¹⁰⁹ J. K. Kraus,²¹ A. Kravchenko,²⁵ S. Kreiss,¹⁰⁹ J. Kretzschmar,⁷³ K. Kreuzfeldt,⁵² N. Krieger,⁵⁴ P. Krieger,¹⁵⁹ K. Kroeninger,⁵⁴ H. Kroha,¹⁰⁰ J. Kroll,¹²¹ J. Kroseberg,²¹ J. Krstic,^{13a} U. Kruchonak,⁶⁴ H. Krüger,²¹ T. Kruker,¹⁷ N. Krumnack,⁶³ Z. V. Krumshateyn,⁶⁴ A. Kruse,¹⁷⁴ M. C. Kruse,⁴⁵ M. Kruskal,²² T. Kubota,⁸⁷ S. Kудay,^{4a} S. Kuehn,⁴⁸ A. Kugel,^{58c} T. Kuhl,⁴² V. Kukhtin,⁶⁴ Y. Kulchitsky,⁹¹ S. Kuleshov,^{32b} M. Kuna,^{133a,133b} J. Kunkle,¹²¹ A. Kupco,¹²⁶ H. Kurashige,⁶⁶ M. Kurata,¹⁶¹ Y. A. Kurochkin,⁹¹ R. Kurumida,⁶⁶ V. Kus,¹²⁶ E. S. Kuwertz,¹⁴⁸ M. Kuze,¹⁵⁸ J. Kvita,¹⁴³ R. Kwee,¹⁶ A. La Rosa,⁴⁹ L. La Rotonda,^{37a,37b} L. Labarga,⁸¹ S. Lablak,^{136a} C. Lacasta,¹⁶⁸ F. Lacava,^{133a,133b} J. Lacey,²⁹ H. Lacker,¹⁶ D. Lacour,⁷⁹ V. R. Lacuesta,¹⁶⁸ E. Ladygin,⁶⁴ R. Lafaye,⁵ B. Laforge,⁷⁹ T. Lagouri,¹⁷⁷ S. Lai,⁴⁸ H. Laier,^{58a} E. Laisne,⁵⁵ L. Lambourne,⁷⁷ C. L. Lampen,⁷ W. Lampl,⁷ E. Lançon,¹³⁷ U. Landgraf,⁴⁸ M. P. J. Landon,⁷⁵ V. S. Lang,^{58a} C. Lange,⁴² A. J. Lankford,¹⁶⁴ F. Lanni,²⁵ K. Lantzsich,³⁰ A. Lanza,^{120a} S. Laplace,⁷⁹ C. Lapoire,²¹ J. F. Laporte,¹³⁷ T. Lari,^{90a} A. Lerner,¹¹⁹ M. Lassnig,³⁰ P. Laurelli,⁴⁷ V. Lavorini,^{37a,37b} W. Lavrijsen,¹⁵ P. Laycock,⁷³ B. T. Le,⁵⁵ O. Le Dortz,⁷⁹ E. Le Guirriec,⁸⁴ E. Le Menedeu,¹² T. LeCompte,⁶ F. Ledroit-Guillon,⁵⁵ C. A. Lee,¹⁵² H. Lee,¹⁰⁶ J. S. H. Lee,¹¹⁷ S. C. Lee,¹⁵² L. Lee,¹⁷⁷ G. Lefebvre,⁷⁹ M. Lefebvre,¹⁷⁰ M. Legendre,¹³⁷ F. Legger,⁹⁹ C. Leggett,¹⁵ A. Lehan,⁷³ M. Lehmacher,²¹ G. Lehmann Miotto,³⁰ A. G. Leister,¹⁷⁷ M. A. L. Leite,^{24d} R. Leitner,¹²⁸ D. Lellouch,¹⁷³ B. Lemmer,⁵⁴ V. Lendermann,^{58a} K. J. C. Leney,^{146c} T. Lenz,¹⁰⁶ G. Lenzen,¹⁷⁶ B. Lenzi,³⁰ R. Leone,⁷ K. Leonhardt,⁴⁴ S. Leontsinis,¹⁰ C. Leroy,⁹⁴ J-R. Lessard,¹⁷⁰ C. G. Lester,²⁸ C. M. Lester,¹²¹ J. Levêque,⁵ D. Levin,⁸⁸ L. J. Levinson,¹⁷³ A. Lewis,¹¹⁹ G. H. Lewis,¹⁰⁹ A. M. Leyko,²¹ M. Leyton,¹⁶ B. Li,^{33b,x} B. Li,⁸⁴ H. Li,¹⁴⁹ H. L. Li,³¹ S. Li,⁴⁵ X. Li,⁸⁸ Z. Liang,^{119,y} H. Liao,³⁴ B. Liberti,^{134a} P. Lichard,³⁰ K. Lie,¹⁶⁶ J. Liebal,²¹ W. Liebig,¹⁴ C. Limbach,²¹ A. Limosani,⁸⁷ M. Limper,⁶² S. C. Lin,^{152,z} F. Linde,¹⁰⁶ B. E. Lindquist,¹⁴⁹ J. T. Linnemann,⁸⁹ E. Lipeles,¹²¹ A. Lipniacka,¹⁴ M. Lisovyi,⁴² T. M. Liss,¹⁶⁶ D. Lissauer,²⁵ A. Lister,¹⁶⁹ A. M. Litke,¹³⁸ B. Liu,¹⁵² D. Liu,¹⁵² J. B. Liu,^{33b} K. Liu,^{33b,aa} L. Liu,⁸⁸ M. Liu,⁴⁵ M. Liu,^{33b} Y. Liu,^{33b} M. Livan,^{120a,120b} S. S. A. Livermore,¹¹⁹ A. Lleres,⁵⁵ J. Llorente Merino,⁸¹ S. L. Lloyd,⁷⁵ F. Lo Sterzo,^{133a,133b} E. Lobodzinska,⁴² P. Loch,⁷ W. S. Lockman,¹³⁸ T. Loddenkoetter,²¹ F. K. Loebinger,⁸³ A. E. Loevschall-Jensen,³⁶ A. Loginov,¹⁷⁷ C. W. Loh,¹⁶⁹ T. Lohse,¹⁶ K. Lohwasser,⁴⁸ M. Lokajicek,¹²⁶ V. P. Lombardo,⁵ J. D. Long,⁸⁸ R. E. Long,⁷¹ L. Lopes,^{125a} D. Lopez Mateos,⁵⁷ B. Lopez Paredes,¹⁴⁰ J. Lorenz,⁹⁹ N. Lorenzo Martinez,¹¹⁶ M. Losada,¹⁶³ P. Loscutoff,¹⁵ M. J. Losty,^{160a,a} X. Lou,⁴¹ A. Lounis,¹¹⁶ J. Love,⁶ P. A. Love,⁷¹ A. J. Lowe,^{144,h} F. Lu,^{33a} H. J. Lubatti,¹³⁹ C. Luci,^{133a,133b} A. Lucotte,⁵⁵ D. Ludwig,⁴² I. Ludwig,⁴⁸ F. Luehring,⁶⁰ W. Lukas,⁶¹ L. Luminari,^{133a} E. Lund,¹¹⁸ J. Lundberg,^{147a,147b} O. Lundberg,^{147a,147b} B. Lund-Jensen,¹⁴⁸ M. Lungwitz,⁸² D. Lynn,²⁵ R. Lysak,¹²⁶ E. Lytken,⁸⁰ H. Ma,²⁵ L. L. Ma,^{33d} G. Maccarrone,⁴⁷ A. Macchiolo,¹⁰⁰ B. Maček,⁷⁴ J. Machado Miguens,^{125a} D. Macina,³⁰ R. Mackeprang,³⁶ R. Madar,⁴⁸ R. J. Madaras,¹⁵ H. J. Maddocks,⁷¹ W. F. Mader,⁴⁴ A. Madsen,¹⁶⁷ M. Maeno,⁸ T. Maeno,²⁵ L. Magnoni,¹⁶⁴ E. Magradze,⁵⁴ K. Mahboubi,⁴⁸ J. Mahlstedt,¹⁰⁶ S. Mahmoud,⁷³ G. Mahout,¹⁸ C. Maiani,¹³⁷ C. Maidantchik,^{24a} A. Maio,^{125a,e} S. Majewski,¹¹⁵ Y. Makida,⁶⁵ N. Makovec,¹¹⁶ P. Mal,^{137,bb} B. Malaescu,⁷⁹ Pa. Malecki,³⁹ V. P. Maleev,¹²² F. Malek,⁵⁵ U. Mallik,⁶² D. Malon,⁶ C. Malone,¹⁴⁴ S. Maltezos,¹⁰ V. M. Malyshev,¹⁰⁸ S. Malyukov,³⁰ J. Mamuzic,^{13b} L. Mandelli,^{90a} I. Mandić,⁷⁴ R. Mandrysch,⁶² J. Maneira,^{125a} A. Manfredini,¹⁰⁰ L. Manhaes de Andrade Filho,^{24b} J. A. Manjarres Ramos,¹³⁷ A. Mann,⁹⁹ P. M. Manning,¹³⁸ A. Manousakis-Katsikakis,⁹ B. Mansoulie,¹³⁷ R. Mantifel,⁸⁶ L. Mapelli,³⁰ L. March,¹⁶⁸ J. F. Marchand,²⁹ F. Marchese,^{134a,134b} G. Marchiori,⁷⁹ M. Marcisovsky,¹²⁶ C. P. Marino,¹⁷⁰ C. N. Marques,^{125a} F. Marroquim,^{24a} Z. Marshall,¹⁵ L. F. Marti,¹⁷ S. Marti-Garcia,¹⁶⁸ B. Martin,³⁰ B. Martin,⁸⁹ J. P. Martin,⁹⁴ T. A. Martin,¹⁷¹ V. J. Martin,⁴⁶ B. Martin dit Latour,⁴⁹ H. Martinez,¹³⁷ M. Martinez,^{12,s} S. Martin-Haugh,¹⁵⁰ A. C. Martyniuk,¹⁷⁰ M. Marx,¹³⁹ F. Marzano,^{133a} A. Marzin,¹¹² L. Masetti,⁸² T. Mashimo,¹⁵⁶ R. Mashinistov,⁹⁵ J. Masik,⁸³ A. L. Maslennikov,¹⁰⁸ I. Massa,^{20a,20b} N. Massol,⁵ P. Mastrandrea,¹⁴⁹ A. Mastroberardino,^{37a,37b} T. Masubuchi,¹⁵⁶ H. Matsunaga,¹⁵⁶ T. Matsushita,⁶⁶ P. Mättig,¹⁷⁶ S. Mättig,⁴² J. Mattmann,⁸² C. Mattravers,^{119,f} J. Maurer,⁸⁴ S. J. Maxfield,⁷³ D. A. Maximov,^{108,i} R. Mazini,¹⁵² L. Mazzaferro,^{134a,134b} M. Mazzanti,^{90a} G. Mc Goldrick,¹⁵⁹ S. P. Mc Kee,⁸⁸ A. McCarn,⁸⁸ R. L. McCarthy,¹⁴⁹ T. G. McCarthy,²⁹ N. A. McCubbin,¹³⁰ K. W. McFarlane,^{56,a} J. A. Mcfayden,¹⁴⁰ G. Mchedlidze,^{51b} T. McLaughlan,¹⁸ S. J. McMahon,¹³⁰ R. A. McPherson,^{170,l} A. Meade,⁸⁵ J. Mechnich,¹⁰⁶ M. Mechtel,¹⁷⁶ M. Medinnis,⁴² S. Meehan,³¹ R. Meera-Lebbai,¹¹² S. Mehlhase,³⁶ A. Mehta,⁷³ K. Meier,^{58a} C. Meineck,⁹⁹ B. Meirose,⁸⁰ C. Melachrinou,³¹ B. R. Mellado Garcia,^{146c} F. Meloni,^{90a,90b} L. Mendoza Navas,¹⁶³ A. Mengarelli,^{20a,20b} S. Menke,¹⁰⁰ E. Meoni,¹⁶² K. M. Mercurio,⁵⁷ S. Mergelmeyer,²¹ N. Meric,¹³⁷ P. Mermod,⁴⁹ L. Merola,^{103a,103b} C. Meroni,^{90a} F. S. Merritt,³¹ H. Merritt,¹¹⁰ A. Messina,^{30,cc} J. Metcalfe,²⁵ A. S. Mete,¹⁶⁴ C. Meyer,⁸² C. Meyer,³¹ J-P. Meyer,¹³⁷ J. Meyer,³⁰ J. Meyer,⁵⁴ S. Michal,³⁰ R. P. Middleton,¹³⁰ S. Migas,⁷³ L. Mijović,¹³⁷ G. Mikenberg,¹⁷³ M. Mikestikova,¹²⁶ M. Mikuž,⁷⁴ D. W. Miller,³¹

W. J. Mills,¹⁶⁹ C. Mills,⁵⁷ A. Milov,¹⁷³ D. A. Milstead,^{147a,147b} D. Milstein,¹⁷³ A. A. Minaenko,¹²⁹
M. Miñano Moya,¹⁶⁸ I. A. Minashvili,⁶⁴ A. I. Mincer,¹⁰⁹ B. Mindur,^{38a} M. Mineev,⁶⁴ Y. Ming,¹⁷⁴ L. M. Mir,¹²
G. Mirabelli,^{133a} T. Mitani,¹⁷² J. Mitrevski,¹³⁸ V. A. Mitsou,¹⁶⁸ S. Mitsui,⁶⁵ P. S. Miyagawa,¹⁴⁰ J. U. Mjörnmark,⁸⁰
T. Moa,^{147a,147b} V. Moeller,²⁸ S. Mohapatra,¹⁴⁹ W. Mohr,⁴⁸ S. Molander,^{147a,147b} R. Moles-Valls,¹⁶⁸ A. Molfetas,³⁰
K. Mönig,⁴² C. Monini,⁵⁵ J. Monk,³⁶ E. Monnier,⁸⁴ J. Montejo Berlingen,¹² F. Monticelli,⁷⁰ S. Monzani,^{20a,20b}
R. W. Moore,³ C. Mora Herrera,⁴⁹ A. Moraes,⁵³ N. Morange,⁶² J. Morel,⁵⁴ D. Moreno,⁸² M. Moreno Llácer,¹⁶⁸
P. Morettini,^{50a} M. Morgenstern,⁴⁴ M. Morii,⁵⁷ S. Moritz,⁸² A. K. Morley,¹⁴⁸ G. Mornacchi,³⁰ J. D. Morris,⁷⁵
L. Morvaj,¹⁰² H. G. Moser,¹⁰⁰ M. Mosidze,^{51b} J. Moss,¹¹⁰ R. Mount,¹⁴⁴ E. Mountricha,^{10,dd} S. V. Mouraviev,^{95,a}
E. J. W. Moyses,⁸⁵ R. D. Mudd,¹⁸ F. Mueller,^{58a} J. Mueller,¹²⁴ K. Mueller,²¹ T. Mueller,²⁸ T. Mueller,⁸²
D. Muenstermann,⁴⁹ Y. Munwes,¹⁵⁴ J. A. Murillo Quijada,¹⁸ W. J. Murray,¹³⁰ I. Mussche,¹⁰⁶ E. Musto,¹⁵³
A. G. Myagkov,^{129,ee} M. Myska,¹²⁶ O. Nackenhorst,⁵⁴ J. Nadal,¹² K. Nagai,⁶¹ R. Nagai,¹⁵⁸ Y. Nagai,⁸⁴ K. Nagano,⁶⁵
A. Nagarkar,¹¹⁰ Y. Nagasaka,⁵⁹ M. Nagel,¹⁰⁰ A. M. Nairz,³⁰ Y. Nakahama,³⁰ K. Nakamura,⁶⁵ T. Nakamura,¹⁵⁶
I. Nakano,¹¹¹ H. Namasivayam,⁴¹ G. Nanava,²¹ A. Napier,¹⁶² R. Narayan,^{58b} M. Nash,^{77,f} T. Nattermann,²¹
T. Naumann,⁴² G. Navarro,¹⁶³ H. A. Neal,⁸⁸ P. Yu. Nechaeva,⁹⁵ T. J. Neep,⁸³ A. Negri,^{120a,120b} G. Negri,³⁰
M. Negrini,^{20a} S. Nektarijevic,⁴⁹ A. Nelson,¹⁶⁴ T. K. Nelson,¹⁴⁴ S. Nemecek,¹²⁶ P. Nemethy,¹⁰⁹
A. A. Nepomuceno,^{24a} M. Nessi,^{30,ff} M. S. Neubauer,¹⁶⁶ M. Neumann,¹⁷⁶ A. Neusiedl,⁸² R. M. Neves,¹⁰⁹ P. Nevski,²⁵
F. M. Newcomer,¹²¹ P. R. Newman,¹⁸ D. H. Nguyen,⁶ V. Nguyen Thi Hong,¹³⁷ R. B. Nickerson,¹¹⁹ R. Nicolaidou,¹³⁷
B. Nicquevert,³⁰ J. Nielsen,¹³⁸ N. Nikiforou,³⁵ A. Nikiforov,¹⁶ V. Nikolaenko,^{129,ee} I. Nikolic-Audit,⁷⁹ K. Nikolics,⁴⁹
K. Nikolopoulos,¹⁸ P. Nilsson,⁸ Y. Ninomiya,¹⁵⁶ A. Nisati,^{133a} R. Nisius,¹⁰⁰ T. Nobe,¹⁵⁸ L. Nodulman,⁶
M. Nomachi,¹¹⁷ I. Nomidis,¹⁵⁵ S. Norberg,¹¹² M. Nordberg,³⁰ J. Novakova,¹²⁸ M. Nozaki,⁶⁵ L. Nozka,¹¹⁴
K. Ntekas,¹⁰ A.-E. Nuncio-Quiroz,²¹ G. Nunes Hanninger,⁸⁷ T. Nunnemann,⁹⁹ E. Nurse,⁷⁷ B. J. O'Brien,⁴⁶
F. O'grady,⁷ D. C. O'Neil,¹⁴³ V. O'Shea,⁵³ L. B. Oakes,⁹⁹ F. G. Oakham,^{29,g} H. Oberlack,¹⁰⁰ J. Ocariz,⁷⁹ A. Ochi,⁶⁶
M. I. Ochoa,⁷⁷ S. Oda,⁶⁹ S. Odaka,⁶⁵ H. Ogren,⁶⁰ A. Oh,⁸³ S. H. Oh,⁴⁵ C. C. Ohm,³⁰ T. Ohshima,¹⁰² W. Okamura,¹¹⁷
H. Okawa,²⁵ Y. Okumura,³¹ T. Okuyama,¹⁵⁶ A. Olariu,^{26a} A. G. Olchevski,⁶⁴ S. A. Olivares Pino,⁴⁶ M. Oliveira,^{125a,j}
D. Oliveira Damazio,²⁵ E. Oliver Garcia,¹⁶⁸ D. Olivito,¹²¹ A. Olszewski,³⁹ J. Olszowska,³⁹ A. Onofre,^{125a,gg}
P. U. E. Onyisi,^{31,hh} C. J. Oram,^{160a} M. J. Oreglia,³¹ Y. Oren,¹⁵⁴ D. Orestano,^{135a,135b} N. Orlando,^{72a,72b}
C. Oropeza Barrera,⁵³ R. S. Orr,¹⁵⁹ B. Osculati,^{50a,50b} R. Ospanov,¹²¹ G. Otero y Garzon,²⁷ H. Otono,⁶⁹
M. Ouchrif,^{136d} E. A. Ouellette,¹⁷⁰ F. Ould-Saada,¹¹⁸ A. Ouraou,¹³⁷ K. P. Oussoren,¹⁰⁶ Q. Ouyang,^{33a}
A. Ovcharova,¹⁵ M. Owen,⁸³ S. Owen,¹⁴⁰ V. E. Ozcan,^{19a} N. Ozturk,⁸ K. Pachal,¹¹⁹ A. Pacheco Pages,¹²
C. Padilla Aranda,¹² S. Pagan Griso,¹⁵ E. Paganis,¹⁴⁰ C. Pahl,¹⁰⁰ F. Paige,²⁵ P. Pais,⁸⁵ K. Pajchel,¹¹⁸ G. Palacino,^{160b}
S. Palestini,³⁰ D. Pallin,³⁴ A. Palma,^{125a} J. D. Palmer,¹⁸ Y. B. Pan,¹⁷⁴ E. Panagiotopoulou,¹⁰ J. G. Panduro Vazquez,⁷⁶
P. Pani,¹⁰⁶ N. Panikashvili,⁸⁸ S. Panitkin,²⁵ D. Pantea,^{26a} Th. D. Papadopoulou,¹⁰ K. Papageorgiou,^{155,r}
A. Paramonov,⁶ D. Paredes Hernandez,³⁴ M. A. Parker,²⁸ F. Parodi,^{50a,50b} J. A. Parsons,³⁵ U. Parzefall,⁴⁸
S. Pashapour,⁵⁴ E. Pasqualucci,^{133a} S. Passaggio,^{50a} A. Passeri,^{135a} F. Pastore,^{135a,135b,a} Fr. Pastore,⁷⁶ G. Pásztor,^{49,ii}
S. Pataraiia,¹⁷⁶ N. D. Patel,¹⁵¹ J. R. Pater,⁸³ S. Patricelli,^{103a,103b} T. Pauly,³⁰ J. Pearce,¹⁷⁰ M. Pedersen,¹¹⁸
S. Pedraza Lopez,¹⁶⁸ M. I. Pedraza Morales,¹⁷⁴ S. V. Peleganchuk,¹⁰⁸ D. Pelikan,¹⁶⁷ H. Peng,^{33b} B. Penning,³¹
A. Penson,³⁵ J. Penwell,⁶⁰ D. V. Perepelitsa,³⁵ T. Perez Cavalcanti,⁴² E. Perez Codina,^{160a}
M. T. Pérez García-Estañ,¹⁶⁸ V. Perez Reale,³⁵ L. Perini,^{90a,90b} H. Pernegger,³⁰ R. Perrino,^{72a} V. D. Peshekhonov,⁶⁴
K. Peters,³⁰ R. F. Y. Peters,^{54,jj} B. A. Petersen,³⁰ J. Petersen,³⁰ T. C. Petersen,³⁶ E. Petit,⁵ A. Petridis,^{147a,147b}
C. Petridou,¹⁵⁵ E. Petrolo,^{133a} F. Petrucci,^{135a,135b} M. Petteni,¹⁴³ R. Pezoa,^{32b} P. W. Phillips,¹³⁰ G. Piacquadio,¹⁴⁴
E. Pianori,¹⁷¹ A. Picazio,⁴⁹ E. Piccaro,⁷⁵ M. Piccinini,^{20a,20b} S. M. Piec,⁴² R. Piegaia,²⁷ D. T. Pignotti,¹¹⁰
J. E. Pilcher,³¹ A. D. Pilkington,⁷⁷ J. Pina,^{125a,e} M. Pinamonti,^{165a,165c,kk} A. Pinder,¹¹⁹ J. L. Pinfold,³ A. Pingel,³⁶
B. Pinto,^{125a} C. Pizio,^{90a,90b} M.-A. Pleier,²⁵ V. Pleskot,¹²⁸ E. Plotnikova,⁶⁴ P. Plucinski,^{147a,147b} S. Poddar,^{58a}
F. Podlyski,³⁴ R. Poettgen,⁸² L. Poggioli,¹¹⁶ D. Pohl,²¹ M. Pohl,⁴⁹ G. Polesello,^{120a} A. Policicchio,^{37a,37b}
R. Polifka,¹⁵⁹ A. Polini,^{20a} C. S. Pollard,⁴⁵ V. Polychronakos,²⁵ D. Pomeroy,²³ K. Pommès,³⁰ L. Pontecorvo,^{133a}
B. G. Pope,⁸⁹ G. A. Popeneciu,^{26b} D. S. Popovic,^{13a} A. Poppleton,³⁰ X. Portell Bueso,¹² G. E. Pospelov,¹⁰⁰
S. Pospisil,¹²⁷ K. Potamianos,¹⁵ I. N. Potrap,⁶⁴ C. J. Potter,¹⁵⁰ C. T. Potter,¹¹⁵ G. Poulard,³⁰ J. Poveda,⁶⁰
V. Pozdnyakov,⁶⁴ R. Prabhu,⁷⁷ P. Pralavorio,⁸⁴ A. Pranko,¹⁵ S. Prasad,³⁰ R. Pravahan,⁸ S. Prell,⁶³ D. Price,⁶⁰
J. Price,⁷³ L. E. Price,⁶ D. Prieur,¹²⁴ M. Primavera,^{72a} M. Proissl,⁴⁶ K. Prokofiev,¹⁰⁹ F. Prokoshin,^{32b}
E. Protopapadaki,¹³⁷ S. Protopopescu,²⁵ J. Proudfoot,⁶ X. Prudent,⁴⁴ M. Przybycien,^{38a} H. Przysieszniak,⁵
S. Psoroulas,²¹ E. Ptacek,¹¹⁵ E. Pueschel,⁸⁵ D. Puldon,¹⁴⁹ M. Purohit,^{25,ll} P. Puzo,¹¹⁶ Y. Pylypchenko,⁶² J. Qian,⁸⁸

- A. Quadt,⁵⁴ D. R. Quarrie,¹⁵ W. B. Quayle,^{146c} D. Quilty,⁵³ V. Radeka,²⁵ V. Radescu,⁴² P. Radloff,¹¹⁵ F. Ragusa,^{90a,90b} G. Rahal,¹⁷⁹ S. Rajagopalan,²⁵ M. Rammensee,⁴⁸ M. Rammes,¹⁴² A. S. Randle-Conde,⁴⁰ C. Rangel-Smith,⁷⁹ K. Rao,¹⁶⁴ F. Rauscher,⁹⁹ T. C. Rave,⁴⁸ T. Ravenscroft,⁵³ M. Raymond,³⁰ A. L. Read,¹¹⁸ D. M. Rebutzi,^{120a,120b} A. Redelbach,¹⁷⁵ G. Redlinger,²⁵ R. Reece,¹²¹ K. Reeves,⁴¹ A. Reinsch,¹¹⁵ I. Reisinger,⁴³ M. Relich,¹⁶⁴ C. Rembser,³⁰ Z. L. Ren,¹⁵² A. Renaud,¹¹⁶ M. Rescigno,^{133a} S. Resconi,^{90a} B. Resende,¹³⁷ P. Reznicek,⁹⁹ R. Rezvani,⁹⁴ R. Richter,¹⁰⁰ E. Richter-Was,^{38b} M. Ridel,⁷⁹ P. Rieck,¹⁶ M. Rijssenbeek,¹⁴⁹ A. Rimoldi,^{120a,120b} L. Rinaldi,^{20a} R. R. Rios,⁴⁰ E. Ritsch,⁶¹ I. Riu,¹² G. Rivoltella,^{90a,90b} F. Rizatdinova,¹¹³ E. Rizvi,⁷⁵ S. H. Robertson,^{86,1} A. Robichaud-Veronneau,¹¹⁹ D. Robinson,²⁸ J. E. M. Robinson,⁸³ A. Robson,⁵³ J. G. Rocha de Lima,¹⁰⁷ C. Roda,^{123a,123b} D. Roda Dos Santos,¹²⁶ L. Rodrigues,³⁰ A. Roe,⁵⁴ S. Roe,³⁰ O. Røhne,¹¹⁸ S. Rolli,¹⁶² A. Romaniouk,⁹⁷ M. Romano,^{20a,20b} G. Romeo,²⁷ E. Romero Adam,¹⁶⁸ N. Rompotis,¹³⁹ L. Roos,⁷⁹ E. Ros,¹⁶⁸ S. Rosati,^{133a} K. Rosbach,⁴⁹ A. Rose,¹⁵⁰ M. Rose,⁷⁶ P. L. Rosendahl,¹⁴ O. Rosenthal,¹⁴² V. Rossetti,¹² E. Rossi,^{103a,103b} L. P. Rossi,^{50a} R. Rosten,¹³⁹ M. Rotaru,^{26a} I. Roth,¹⁷³ J. Rothberg,¹³⁹ D. Rousseau,¹¹⁶ C. R. Royon,¹³⁷ A. Rozanov,⁸⁴ Y. Rozen,¹⁵³ X. Ruan,^{146c} F. Rubbo,¹² I. Rubinskiy,⁴² V. I. Rud,⁹⁸ C. Rudolph,⁴⁴ M. S. Rudolph,¹⁵⁹ F. Rühr,⁷ A. Ruiz-Martinez,⁶³ L. Rummyantsev,⁶⁴ Z. Rurikova,⁴⁸ N. A. Rusakovich,⁶⁴ A. Ruschke,⁹⁹ J. P. Rutherford,⁷ N. Ruthmann,⁴⁸ P. Ruzicka,¹²⁶ Y. F. Ryabov,¹²² M. Rybar,¹²⁸ G. Rybkin,¹¹⁶ N. C. Ryder,¹¹⁹ A. F. Saavedra,¹⁵¹ A. Saddique,³ I. Sadeh,¹⁵⁴ H. F. W. Sadrozinski,¹³⁸ R. Sadykov,⁶⁴ F. Safai Tehrani,^{133a} H. Sakamoto,¹⁵⁶ Y. Sakurai,¹⁷² G. Salamanna,⁷⁵ A. Salamon,^{134a} M. Saleem,¹¹² D. Salek,³⁰ D. Salihagic,¹⁰⁰ A. Salnikov,¹⁴⁴ J. Salt,¹⁶⁸ B. M. Salvachua Ferrando,⁶ D. Salvatore,^{37a,37b} F. Salvatore,¹⁵⁰ A. Salvucci,¹⁰⁵ A. Salzburger,³⁰ D. Sampsonidis,¹⁵⁵ A. Sanchez,^{103a,103b} J. Sánchez,¹⁶⁸ V. Sanchez Martinez,¹⁶⁸ H. Sandaker,¹⁴ H. G. Sander,⁸² M. P. Sanders,⁹⁹ M. Sandhoff,¹⁷⁶ T. Sandoval,²⁸ C. Sandoval,¹⁶³ R. Sandstroem,¹⁰⁰ D. P. C. Sankey,¹³⁰ A. Sansoni,⁴⁷ C. Santoni,³⁴ R. Santonico,^{134a,134b} H. Santos,^{125a} I. Santoyo Castillo,¹⁵⁰ K. Sapp,¹²⁴ A. Saproinov,⁶⁴ J. G. Saraiva,^{125a} E. Sarkisyan-Grinbaum,⁸ B. Sarrazin,²¹ G. Sartisohn,¹⁷⁶ O. Sasaki,⁶⁵ Y. Sasaki,¹⁵⁶ N. Sasao,⁶⁷ I. Satsounkevitch,⁹¹ G. Sauvage,^{5,a} E. Sauvan,⁵ J. B. Sauvan,¹¹⁶ P. Savard,^{159,g} V. Savinov,¹²⁴ D. O. Savu,³⁰ C. Sawyer,¹¹⁹ L. Sawyer,^{78,n} D. H. Saxon,⁵³ J. Saxon,¹²¹ C. Sbarra,^{20a} A. Sbrizzi,³ T. Scanlon,³⁰ D. A. Scannicchio,¹⁶⁴ M. Scarcella,¹⁵¹ J. Schaarschmidt,¹¹⁶ P. Schacht,¹⁰⁰ D. Schaefer,¹²¹ A. Schaelicke,⁴⁶ S. Schaepe,²¹ S. Schaezel,^{58b} U. Schäfer,⁸² A. C. Schaffer,¹¹⁶ D. Schaile,⁹⁹ R. D. Schamberger,¹⁴⁹ V. Scharf,^{58a} V. A. Schegelsky,¹²² D. Scheirich,⁸⁸ M. Schernau,¹⁶⁴ M. I. Scherzer,³⁵ C. Schiavi,^{50a,50b} J. Schieck,⁹⁹ C. Schillo,⁴⁸ M. Schioppa,^{37a,37b} S. Schlenker,³⁰ E. Schmidt,⁴⁸ K. Schmieden,³⁰ C. Schmitt,⁸² C. Schmitt,⁹⁹ S. Schmitt,^{58b} B. Schneider,¹⁷ Y. J. Schnellbach,⁷³ U. Schnoor,⁴⁴ L. Schoeffel,¹³⁷ A. Schoening,^{58b} B. D. Schoenrock,⁸⁹ A. L. S. Schorlemmer,⁵⁴ M. Schott,⁸² D. Schouten,^{160a} J. Schovancova,²⁵ M. Schram,⁸⁶ S. Schramm,¹⁵⁹ M. Schreyer,¹⁷⁵ C. Schroeder,⁸² N. Schroer,^{58c} N. Schuh,⁸² M. J. Schultens,²¹ H.-C. Schultz-Coulon,^{58a} H. Schulz,¹⁶ M. Schumacher,⁴⁸ B. A. Schumm,¹³⁸ Ph. Schune,¹³⁷ A. Schwartzman,¹⁴⁴ Ph. Schwegler,¹⁰⁰ Ph. Schwemling,¹³⁷ R. Schwienhorst,⁸⁹ J. Schwindling,¹³⁷ T. Schwindt,²¹ M. Schwoerer,⁵ F. G. Sciacca,¹⁷ E. Scifo,¹¹⁶ G. Sciolla,²³ W. G. Scott,¹³⁰ F. Scutti,²¹ J. Searcy,⁸⁸ G. Sedov,⁴² E. Sedykh,¹²² S. C. Seidel,¹⁰⁴ A. Seiden,¹³⁸ F. Seifert,⁴⁴ J. M. Seixas,^{24a} G. Sekhniaidze,^{103a} S. J. Sekula,⁴⁰ K. E. Selbach,⁴⁶ D. M. Seliverstov,¹²² G. Sellers,⁷³ M. Seman,^{145b} N. Semprini-Cesari,^{20a,20b} C. Serfon,³⁰ L. Serin,¹¹⁶ L. Serkin,⁵⁴ T. Serre,⁸⁴ R. Seuster,^{160a} H. Severini,¹¹² F. Sforza,¹⁰⁰ A. Sfyrta,³⁰ E. Shabalina,⁵⁴ M. Shamim,¹¹⁵ L. Y. Shan,^{33a} J. T. Shank,²² Q. T. Shao,⁸⁷ M. Shapiro,¹⁵ P. B. Shatalov,⁹⁶ K. Shaw,^{165a,165c} P. Sherwood,⁷⁷ S. Shimizu,⁶⁶ M. Shimojima,¹⁰¹ T. Shin,⁵⁶ M. Shiyakova,⁶⁴ A. Shmeleva,⁹⁵ M. J. Shochet,³¹ D. Short,¹¹⁹ S. Shrestha,⁶³ E. Shulga,⁹⁷ M. A. Shupe,⁷ S. Shushkevich,⁴² P. Sicho,¹²⁶ D. Sidorov,¹¹³ A. Sidoti,^{133a} F. Siegert,⁴⁸ Dj. Sijacki,^{13a} O. Silbert,¹⁷³ J. Silva,^{125a} Y. Silver,¹⁵⁴ D. Silverstein,¹⁴⁴ S. B. Silverstein,^{147a} V. Simak,¹²⁷ O. Simard,⁵ Lj. Simic,^{13a} S. Simion,¹¹⁶ E. Simioni,⁸² B. Simmons,⁷⁷ R. Simoniello,^{90a,90b} M. Simonyan,³⁶ P. Sinervo,¹⁵⁹ N. B. Sinev,¹¹⁵ V. Sipica,¹⁴² G. Siragusa,¹⁷⁵ A. Sircar,⁷⁸ A. N. Sisakyan,^{64,a} S. Yu. Sivoklovov,⁹⁸ J. Sjölin,^{147a,147b} T. B. Sjrursen,¹⁴ L. A. Skinnari,¹⁵ H. P. Skottowe,⁵⁷ K. Yu. Skovpen,¹⁰⁸ P. Skubic,¹¹² M. Slater,¹⁸ T. Slavicek,¹²⁷ K. Sliwa,¹⁶² V. Smakhtin,¹⁷³ B. H. Smart,⁴⁶ L. Smestad,¹¹⁸ S. Yu. Smirnov,⁹⁷ Y. Smirnov,⁹⁷ L. N. Smirnova,^{98,mm} O. Smirnova,⁸⁰ K. M. Smith,⁵³ M. Smizanska,⁷¹ K. Smolek,¹²⁷ A. A. Snesarev,⁹⁵ G. Snidero,⁷⁵ J. Snow,¹¹² S. Snyder,²⁵ R. Sobie,^{170,1} F. Socher,⁴⁴ J. Sodomka,¹²⁷ A. Soffer,¹⁵⁴ D. A. Soh,^{152,y} C. A. Solans,³⁰ M. Solar,¹²⁷ J. Solc,¹²⁷ E. Yu. Soldatov,⁹⁷ U. Soldevila,¹⁶⁸ E. Solfaroli Camillocci,^{133a,133b} A. A. Solodkov,¹²⁹ O. V. Solovyanov,¹²⁹ V. Solovyev,¹²² N. Soni,¹ A. Sood,¹⁵ V. Sopko,¹²⁷ B. Sopko,¹²⁷ M. Sosebee,⁸ R. Soualah,^{165a,165c} P. Soueid,⁹⁴ A. M. Soukharev,¹⁰⁸ D. South,⁴² S. Spagnolo,^{72a,72b} F. Spanò,⁷⁶ W. R. Spearman,⁵⁷ R. Spighi,^{20a} G. Spigo,³⁰ M. Spousta,^{128,nn} T. Spreitzer,¹⁵⁹ B. Spurlock,⁸ R. D. St. Denis,⁵³ J. Stahlman,¹²¹ R. Stamen,^{58a} E. Stanecka,³⁹ R. W. Stanek,⁶ C. Stanescu,^{135a}

- M. Stanescu-Bellu,⁴² M. M. Stanitzki,⁴² S. Stapnes,¹¹⁸ E. A. Starchenko,¹²⁹ J. Stark,⁵⁵ P. Staroba,¹²⁶ P. Starovoitov,⁴² R. Staszewski,³⁹ P. Stavina,^{145a,a} G. Steele,⁵³ P. Steinbach,⁴⁴ P. Steinberg,²⁵ I. Stekl,¹²⁷ B. Stelzer,¹⁴³ H. J. Stelzer,⁸⁹ O. Stelzer-Chilton,^{160a} H. Stenzel,⁵² S. Stern,¹⁰⁰ G. A. Stewart,³⁰ J. A. Stillings,²¹ M. C. Stockton,⁸⁶ M. Stoebe,⁸⁶ K. Stoerig,⁴⁸ G. Stoicea,^{26a} S. Stonjek,¹⁰⁰ A. R. Stradling,⁸ A. Straessner,⁴⁴ J. Strandberg,¹⁴⁸ S. Strandberg,^{147a,147b} A. Strandlie,¹¹⁸ E. Strauss,¹⁴⁴ M. Strauss,¹¹² P. Strizenec,^{145b} R. Ströhmer,¹⁷⁵ D. M. Strom,¹¹⁵ R. Stroynowski,⁴⁰ S. A. Stucci,¹⁷ B. Stugu,¹⁴ I. Stumer,^{25,a} J. Stupak,¹⁴⁹ P. Sturm,¹⁷⁶ N. A. Styles,⁴² D. Su,¹⁴⁴ HS. Subramania,³ R. Subramaniam,⁷⁸ A. Succurro,¹² Y. Sugaya,¹¹⁷ C. Suhr,¹⁰⁷ M. Suk,¹²⁷ V. V. Sulin,⁹⁵ S. Sultansoy,^{4c} T. Sumida,⁶⁷ X. Sun,⁵⁵ J. E. Sundermann,⁴⁸ K. Suruliz,¹⁴⁰ G. Susinno,^{37a,37b} M. R. Sutton,¹⁵⁰ Y. Suzuki,⁶⁵ M. Svatos,¹²⁶ S. Swedish,¹⁶⁹ M. Swiatlowski,¹⁴⁴ I. Sykora,^{145a} T. Sykora,¹²⁸ D. Ta,⁸⁹ K. Tackmann,⁴² J. Taenzer,¹⁵⁹ A. Taffard,¹⁶⁴ R. Tahirout,^{160a} N. Taiblum,¹⁵⁴ Y. Takahashi,¹⁰² H. Takai,²⁵ R. Takashima,⁶⁸ H. Takeda,⁶⁶ T. Takeshita,¹⁴¹ Y. Takubo,⁶⁵ M. Talby,⁸⁴ A. A. Talyshv,^{108,i} J. Y. C. Tam,¹⁷⁵ M. C. Tamsett,^{78,oo} K. G. Tan,⁸⁷ J. Tanaka,¹⁵⁶ R. Tanaka,¹¹⁶ S. Tanaka,¹³² S. Tanaka,⁶⁵ A. J. Tanasijczuk,¹⁴³ K. Tani,⁶⁶ N. Tannoury,⁸⁴ S. Tapprogge,⁸² S. Tarem,¹⁵³ F. Tarrade,²⁹ G. F. Tartarelli,^{90a} P. Tas,¹²⁸ M. Tasevsky,¹²⁶ T. Tashiro,⁶⁷ E. Tassi,^{37a,37b} A. Tavares Delgado,^{125a} Y. Tayalati,^{136d} C. Taylor,⁷⁷ F. E. Taylor,⁹³ G. N. Taylor,⁸⁷ W. Taylor,^{160b} F. A. Teischinger,³⁰ M. Teixeira Dias Castanheira,⁷⁵ P. Teixeira-Dias,⁷⁶ K. K. Temming,⁴⁸ H. Ten Kate,³⁰ P. K. Teng,¹⁵² S. Terada,⁶⁵ K. Terashi,¹⁵⁶ J. Terron,⁸¹ S. Terzo,¹⁰⁰ M. Testa,⁴⁷ R. J. Teuscher,^{159,1} J. Therhaag,²¹ T. Theveneaux-Pelzer,³⁴ S. Thoma,⁴⁸ J. P. Thomas,¹⁸ E. N. Thompson,³⁵ P. D. Thompson,¹⁸ P. D. Thompson,¹⁵⁹ A. S. Thompson,⁵³ L. A. Thomsen,³⁶ E. Thomson,¹²¹ M. Thomson,²⁸ W. M. Thong,⁸⁷ R. P. Thun,^{88,a} F. Tian,³⁵ M. J. Tibbetts,¹⁵ T. Tic,¹²⁶ V. O. Tikhomirov,^{95,pp} Yu. A. Tikhonov,^{108,i} S. Timoshenko,⁹⁷ E. Tiouchichine,⁸⁴ P. Tipton,¹⁷⁷ S. Tisserant,⁸⁴ T. Todorov,⁵ S. Todorova-Nova,¹²⁸ B. Toggerson,¹⁶⁴ J. Tojo,⁶⁹ S. Tokár,^{145a} K. Tokushuku,⁶⁵ K. Tollefson,⁸⁹ L. Tomlinson,⁸³ M. Tomoto,¹⁰² L. Tompkins,³¹ K. Toms,¹⁰⁴ A. Tonoyan,¹⁴ N. D. Topilin,⁶⁴ E. Torrence,¹¹⁵ H. Torres,¹⁴³ E. Torró Pastor,¹⁶⁸ J. Toth,^{84,ii} F. Touchard,⁸⁴ D. R. Tovey,¹⁴⁰ H. L. Tran,¹¹⁶ T. Trefzger,¹⁷⁵ L. Tremblet,³⁰ A. Tricoli,³⁰ I. M. Trigger,^{160a} S. Trincaz-Duvold,⁷⁹ M. F. Tripiana,⁷⁰ N. Triplett,²⁵ W. Trischuk,¹⁵⁹ B. Trocmé,⁵⁵ C. Troncon,^{90a} M. Trotter-McDonald,¹⁴³ M. Trovatelli,^{135a,135b} P. True,⁸⁹ M. Trzebinski,³⁹ A. Trzupek,³⁹ C. Tsarouchas,³⁰ J. C-L. Tseng,¹¹⁹ P. V. Tsiareshka,⁹¹ D. Tsiouou,¹³⁷ G. Tsipolitis,¹⁰ N. Tsirintanis,⁹ S. Tsiskaridze,¹² V. Tsiskaridze,⁴⁸ E. G. Tskhadadze,^{51a} I. I. Tsukerman,⁹⁶ V. Tsulaia,¹⁵ J.-W. Tsung,²¹ S. Tsuno,⁶⁵ D. Tsybychev,¹⁴⁹ A. Tua,¹⁴⁰ A. Tudorache,^{26a} V. Tudorache,^{26a} J. M. Tuggle,³¹ A. N. Tuna,¹²¹ S. A. Tuppiti,^{20a,20b} S. Turchikhin,^{98,mmm} D. Turecek,¹²⁷ I. Turk Cakir,^{4d} R. Turra,^{90a,90b} P. M. Tuts,³⁵ A. Tykhonov,⁷⁴ M. Tylmad,^{147a,147b} M. Tyndel,¹³⁰ K. Uchida,²¹ I. Ueda,¹⁵⁶ R. Ueno,²⁹ M. Ughetto,⁸⁴ M. Ugland,¹⁴ M. Uhlenbrock,²¹ F. Ukegawa,¹⁶¹ G. Unal,³⁰ A. Undrus,²⁵ G. Unel,¹⁶⁴ F. C. Ungaro,⁴⁸ Y. Unno,⁶⁵ D. Urbaniec,³⁵ P. Urquijo,²¹ G. Usai,⁸ A. Usanova,⁶¹ L. Vacavant,⁸⁴ V. Vacek,¹²⁷ B. Vachon,⁸⁶ S. Vahsen,¹⁵ N. Valencic,¹⁰⁶ S. Valentinetti,^{20a,20b} A. Valero,¹⁶⁸ L. Valery,³⁴ S. Valkar,¹²⁸ E. Valladolid Gallego,¹⁶⁸ S. Vallecorsa,⁴⁹ J. A. Valls Ferrer,¹⁶⁸ R. Van Berg,¹²¹ P. C. Van Der Deijl,¹⁰⁶ R. van der Geer,¹⁰⁶ H. van der Graaf,¹⁰⁶ R. Van Der Leeuw,¹⁰⁶ D. van der Ster,³⁰ N. van Eldik,³⁰ P. van Gemmeren,⁶ J. Van Nieuwkoop,¹⁴³ I. van Vulpen,¹⁰⁶ M. C. van Woerden,³⁰ M. Vanadia,¹⁰⁰ W. Vandelli,³⁰ A. Vaniachine,⁶ P. Vankov,⁴² F. Vannucci,⁷⁹ R. Vari,^{133a} E. W. Varnes,⁷ T. Varol,⁸⁵ D. Varouchas,¹⁵ A. Vartapetian,⁸ K. E. Varvell,¹⁵¹ V. I. Vassilakopoulos,⁵⁶ F. Vazeille,³⁴ T. Vazquez Schroeder,⁵⁴ J. Veatch,⁷ F. Veloso,^{125a} S. Veneziano,^{133a} A. Ventura,^{72a,72b} D. Ventura,⁸⁵ M. Venturi,⁴⁸ N. Venturi,¹⁵⁹ V. Vercesi,^{120a} M. Verducci,¹³⁹ W. Verkerke,¹⁰⁶ J. C. Vermeulen,¹⁰⁶ A. Vest,⁴⁴ M. C. Vetterli,^{143,g} O. Viazlo,⁸⁰ I. Vichou,¹⁶⁶ T. Vickey,^{146c,qq} O. E. Vickey Boeriu,^{146c} G. H. A. Viehhauser,¹¹⁹ S. Viel,¹⁶⁹ R. Vigne,³⁰ M. Villa,^{20a,20b} M. Villaplana Perez,¹⁶⁸ E. Vilucchi,⁴⁷ M. G. Vinciter,²⁹ V. B. Vinogradov,⁶⁴ J. Virzi,¹⁵ O. Vitells,¹⁷³ M. Viti,⁴² I. Vivarelli,¹⁵⁰ F. Vives Vaque,³ S. Vlachos,¹⁰ D. Vladoiu,⁹⁹ M. Vlasak,¹²⁷ A. Vogel,²¹ P. Vokac,¹²⁷ G. Volpi,⁴⁷ M. Volpi,⁸⁷ G. Volpini,^{90a} H. von der Schmitt,¹⁰⁰ H. von Radziewski,⁴⁸ E. von Toerne,²¹ V. Vorobel,¹²⁸ M. Vos,¹⁶⁸ R. Voss,³⁰ J. H. Vosseveld,⁷³ N. Vranjes,¹³⁷ M. Vranjes Milosavljevic,¹⁰⁶ V. Vrba,¹²⁶ M. Vreeswijk,¹⁰⁶ T. Vu Anh,⁴⁸ R. Vuillermet,³⁰ I. Vukotic,³¹ Z. Vykydal,¹²⁷ W. Wagner,¹⁷⁶ P. Wagner,²¹ S. Wahrmund,⁴⁴ J. Wakabayashi,¹⁰² S. Walch,⁸⁸ J. Walder,⁷¹ R. Walker,⁹⁹ W. Walkowiak,¹⁴² R. Wall,¹⁷⁷ P. Waller,⁷³ B. Walsh,¹⁷⁷ C. Wang,⁴⁵ H. Wang,¹⁷⁴ H. Wang,⁴⁰ J. Wang,¹⁵² J. Wang,^{33a} K. Wang,⁸⁶ R. Wang,¹⁰⁴ S. M. Wang,¹⁵² T. Wang,²¹ X. Wang,¹⁷⁷ A. Warburton,⁸⁶ C. P. Ward,²⁸ D. R. Wardrope,⁷⁷ M. Warsinsky,⁴⁸ A. Washbrook,⁴⁶ C. Wasicki,⁴² I. Watanabe,⁶⁶ P. M. Watkins,¹⁸ A. T. Watson,¹⁸ I. J. Watson,¹⁵¹ M. F. Watson,¹⁸ G. Watts,¹³⁹ S. Watts,⁸³ A. T. Waugh,¹⁵¹ B. M. Waugh,⁷⁷ S. Webb,⁸³ M. S. Weber,¹⁷ S. W. Weber,¹⁷⁵ J. S. Webster,³¹ A. R. Weidberg,¹¹⁹ P. Weigell,¹⁰⁰ J. Weingarten,⁵⁴ C. Weiser,⁴⁸ H. Weits,¹⁰⁶ P. S. Wells,³⁰ T. Wenaus,²⁵ D. Wendland,¹⁶ Z. Weng,^{152,y} T. Wengler,³⁰ S. Wenig,³⁰ N. Wermes,²¹ M. Werner,⁴⁸ P. Werner,³⁰ M. Wessels,^{58a} J. Wetter,¹⁶² K. Whalen,²⁹ A. White,⁸ M. J. White,¹ R. White,^{32b} S. White,^{123a,123b} D. Whiteson,¹⁶⁴

D. Whittington,⁶⁰ D. Wicke,¹⁷⁶ F. J. Wickens,¹³⁰ W. Wiedenmann,¹⁷⁴ M. Wielers,^{80,f} P. Wienemann,²¹
 C. Wiglesworth,³⁶ L. A. M. Wiik-Fuchs,²¹ P. A. Wijeratne,⁷⁷ A. Wildauer,¹⁰⁰ M. A. Wildt,^{42,tr} I. Wilhelm,¹²⁸
 H. G. Wilkens,³⁰ J. Z. Will,⁹⁹ E. Williams,³⁵ H. H. Williams,¹²¹ S. Williams,²⁸ W. Willis,^{35,a} S. Willocq,⁸⁵
 J. A. Wilson,¹⁸ A. Wilson,⁸⁸ I. Wingerter-Seez,⁵ S. Winkelmann,⁴⁸ F. Winklmeier,¹¹⁵ M. Wittgen,¹⁴⁴ T. Wittig,⁴³
 J. Wittkowski,⁹⁹ S. J. Wollstadt,⁸² M. W. Wolter,³⁹ H. Wolters,^{125a,j} W. C. Wong,⁴¹ B. K. Wosiek,³⁹ J. Wotschack,³⁰
 M. J. Woudstra,⁸³ K. W. Wozniak,³⁹ K. Wraight,⁵³ M. Wright,⁵³ S. L. Wu,¹⁷⁴ X. Wu,⁴⁹ Y. Wu,⁸⁸ E. Wulf,³⁵
 T. R. Wyatt,⁸³ B. M. Wynne,⁴⁶ S. Xella,³⁶ M. Xiao,¹³⁷ C. Xu,^{33b,dd} D. Xu,^{33a} L. Xu,^{33b,ss} B. Yabsley,¹⁵¹
 S. Yacoob,^{146b,tt} M. Yamada,⁶⁵ H. Yamaguchi,¹⁵⁶ Y. Yamaguchi,¹⁵⁶ A. Yamamoto,⁶⁵ K. Yamamoto,⁶³
 S. Yamamoto,¹⁵⁶ T. Yamamura,¹⁵⁶ T. Yamanaka,¹⁵⁶ K. Yamauchi,¹⁰² Y. Yamazaki,⁶⁶ Z. Yan,²² H. Yang,^{33e}
 H. Yang,¹⁷⁴ U. K. Yang,⁸³ Y. Yang,¹¹⁰ Z. Yang,^{147a,147b} S. Yanush,⁹² L. Yao,^{33a} Y. Yasu,⁶⁵ E. Yatsenko,⁴²
 K. H. Yau Wong,²¹ J. Ye,⁴⁰ S. Ye,²⁵ A. L. Yen,⁵⁷ E. Yildirim,⁴² M. Yilmaz,^{4b} R. Yoosofmiya,¹²⁴ K. Yorita,¹⁷²
 R. Yoshida,⁶ K. Yoshihara,¹⁵⁶ C. Young,¹⁴⁴ C. J. S. Young,¹¹⁹ S. Youssef,²² D. R. Yu,¹⁵ J. Yu,⁸ J. Yu,¹¹³ L. Yuan,⁶⁶
 A. Yurkewicz,¹⁰⁷ B. Zabinski,³⁹ R. Zaidan,⁶² A. M. Zaitsev,^{129,ee} S. Zambito,²³ L. Zanello,^{133a,133b} D. Zanzi,¹⁰⁰
 A. Zaytsev,²⁵ C. Zeitnitz,¹⁷⁶ M. Zeman,¹²⁷ A. Zemla,³⁹ O. Zenin,¹²⁹ T. Ženiš,^{145a} D. Zerwas,¹¹⁶ G. Zevi della Porta,⁵⁷
 D. Zhang,⁸⁸ H. Zhang,⁸⁹ J. Zhang,⁶ L. Zhang,¹⁵² X. Zhang,^{33d} Z. Zhang,¹¹⁶ Z. Zhao,^{33b} A. Zhemchugov,⁶⁴
 J. Zhong,¹¹⁹ B. Zhou,⁸⁸ L. Zhou,³⁵ N. Zhou,¹⁶⁴ C. G. Zhu,^{33d} H. Zhu,⁴² J. Zhu,⁸⁸ Y. Zhu,^{33b} X. Zhuang,^{33a} A. Zibell,⁹⁹
 D. Zieminska,⁶⁰ N. I. Zimin,⁶⁴ C. Zimmermann,⁸² R. Zimmermann,²¹ S. Zimmermann,²¹ S. Zimmermann,⁴⁸
 Z. Zinonos,^{123a,123b} M. Ziolkowski,¹⁴² R. Zitoun,⁵ L. Živković,³⁵ G. Zobernig,¹⁷⁴ A. Zoccoli,^{20a,20b} M. zur Nedden,¹⁶
 G. Zurzolo,^{103a,103b} V. Zutshi,¹⁰⁷ and L. Zwalinski³⁰

(ATLAS Collaboration)^{uu}

¹*School of Chemistry and Physics, University of Adelaide, Adelaide, Australia*

²*Physics Department, SUNY Albany, Albany, New York, USA*

³*Department of Physics, University of Alberta, Edmonton, Alberta, Canada*

^{4a}*Department of Physics, Ankara University, Ankara, Turkey*

^{4b}*Department of Physics, Gazi University, Ankara, Turkey*

^{4c}*Division of Physics, TOBB University of Economics and Technology, Ankara, Turkey*

^{4d}*Turkish Atomic Energy Authority, Ankara, Turkey*

⁵*LAPP, CNRS/IN2P3 and Université de Savoie, Annecy-le-Vieux, France*

⁶*High Energy Physics Division, Argonne National Laboratory, Argonne, Illinois, USA*

⁷*Department of Physics, University of Arizona, Tucson, Arizona, USA*

⁸*Department of Physics, The University of Texas at Arlington, Arlington, Texas, USA*

⁹*Physics Department, University of Athens, Athens, Greece*

¹⁰*Physics Department, National Technical University of Athens, Zografou, Greece*

¹¹*Institute of Physics, Azerbaijan Academy of Sciences, Baku, Azerbaijan*

¹²*Institut de Física d'Altes Energies and Departament de Física de la Universitat Autònoma de Barcelona, Barcelona, Spain*

^{13a}*Institute of Physics, University of Belgrade, Belgrade, Serbia*

^{13b}*Vinca Institute of Nuclear Sciences, University of Belgrade, Belgrade, Serbia*

¹⁴*Department for Physics and Technology, University of Bergen, Bergen, Norway*

¹⁵*Physics Division, Lawrence Berkeley National Laboratory and University of California, Berkeley, California, USA*

¹⁶*Department of Physics, Humboldt University, Berlin, Germany*

¹⁷*Albert Einstein Center for Fundamental Physics and Laboratory for High Energy Physics, University of Bern, Bern, Switzerland*

¹⁸*School of Physics and Astronomy, University of Birmingham, Birmingham, United Kingdom*

^{19a}*Department of Physics, Bogazici University, Istanbul, Turkey*

^{19b}*Department of Physics, Dogus University, Istanbul, Turkey*

^{19c}*Department of Physics Engineering, Gaziantep University, Gaziantep, Turkey*

^{20a}*INFN Sezione di Bologna, Italy*

^{20b}*Dipartimento di Fisica e Astronomia, Università di Bologna, Bologna, Italy*

²¹*Physikalisches Institut, University of Bonn, Bonn, Germany*

²²*Department of Physics, Boston University, Boston, Massachusetts*

²³*Department of Physics, Brandeis University, Waltham, Massachusetts, USA*

^{24a}*Universidade Federal do Rio De Janeiro COPPE/EE/IF, Rio de Janeiro, Brazil*

^{24b}*Federal University of Juiz de Fora (UFJF), Juiz de Fora, Brazil*

^{24c}*Federal University of Sao Joao del Rei (UFSJ), Sao Joao del Rei, Brazil*

^{24d}*Instituto de Física, Universidade de Sao Paulo, Sao Paulo, Brazil*

- ²⁵Physics Department, Brookhaven National Laboratory, Upton, New York, USA
- ^{26a}National Institute of Physics and Nuclear Engineering, Bucharest, Romania
- ^{26b}National Institute for Research and Development of Isotopic and Molecular Technologies, Physics Department, Cluj Napoca, Romania
- ^{26c}University Politehnica Bucharest, Bucharest, Romania
- ^{26d}West University in Timisoara, Timisoara, Romania
- ²⁷Departamento de Física, Universidad de Buenos Aires, Buenos Aires, Argentina
- ²⁸Cavendish Laboratory, University of Cambridge, Cambridge, United Kingdom
- ²⁹Department of Physics, Carleton University, Ottawa, Ontario, Canada
- ³⁰CERN, Geneva, Switzerland
- ³¹Enrico Fermi Institute, University of Chicago, Chicago, Illinois, USA
- ^{32a}Departamento de Física, Pontificia Universidad Católica de Chile, Santiago, Chile
- ^{32b}Departamento de Física, Universidad Técnica Federico Santa María, Valparaíso, Chile
- ^{33a}Institute of High Energy Physics, Chinese Academy of Sciences, Beijing, China
- ^{33b}Department of Modern Physics, University of Science and Technology of China, Anhui, China
- ^{33c}Department of Physics, Nanjing University, Jiangsu, China
- ^{33d}School of Physics, Shandong University, Shandong, China
- ^{33e}Physics Department, Shanghai Jiao Tong University, Shanghai, China
- ³⁴Laboratoire de Physique Corpusculaire, Clermont Université and Université Blaise Pascal and CNRS/IN2P3, Clermont-Ferrand, France
- ³⁵Nevis Laboratory, Columbia University, Irvington, New York, USA
- ³⁶Niels Bohr Institute, University of Copenhagen, Kobenhavn, Denmark
- ^{37a}INFN Gruppo Collegato di Cosenza, Italy
- ^{37b}Dipartimento di Fisica, Università della Calabria, Rende, Italy
- ^{38a}AGH University of Science and Technology, Faculty of Physics and Applied Computer Science, Krakow, Poland
- ^{38b}Marian Smoluchowski Institute of Physics, Jagiellonian University, Krakow, Poland
- ³⁹The Henryk Niewodniczanski Institute of Nuclear Physics, Polish Academy of Sciences, Krakow, Poland
- ⁴⁰Physics Department, Southern Methodist University, Dallas, Texas, USA
- ⁴¹Physics Department, University of Texas at Dallas, Richardson, Texas, USA
- ⁴²DESY, Hamburg and Zeuthen, Germany
- ⁴³Institut für Experimentelle Physik IV, Technische Universität Dortmund, Dortmund, Germany
- ⁴⁴Institut für Kern- und Teilchenphysik, Technische Universität Dresden, Dresden, Germany
- ⁴⁵Department of Physics, Duke University, Durham, North Carolina, USA
- ⁴⁶SUPA-School of Physics and Astronomy, University of Edinburgh, Edinburgh, United Kingdom
- ⁴⁷INFN Laboratori Nazionali di Frascati, Frascati, Italy
- ⁴⁸Fakultät für Mathematik und Physik, Albert-Ludwigs-Universität, Freiburg, Germany
- ⁴⁹Section de Physique, Université de Genève, Geneva, Switzerland
- ^{50a}INFN Sezione di Genova, Italy
- ^{50b}Dipartimento di Fisica, Università di Genova, Genova, Italy
- ^{51a}E. Andronikashvili Institute of Physics, Iv. Javakhishvili Tbilisi State University, Tbilisi, Georgia
- ^{51b}High Energy Physics Institute, Tbilisi State University, Tbilisi, Georgia
- ⁵²II Physikalisches Institut, Justus-Liebig-Universität Giessen, Giessen, Germany
- ⁵³SUPA-School of Physics and Astronomy, University of Glasgow, Glasgow, United Kingdom
- ⁵⁴II Physikalisches Institut, Georg-August-Universität, Göttingen, Germany
- ⁵⁵Laboratoire de Physique Subatomique et de Cosmologie, Université Joseph Fourier and CNRS/IN2P3 and Institut National Polytechnique de Grenoble, Grenoble, France
- ⁵⁶Department of Physics, Hampton University, Hampton, Virginia, USA
- ⁵⁷Laboratory for Particle Physics and Cosmology, Harvard University, Cambridge, Massachusetts, USA
- ^{58a}Kirchhoff-Institut für Physik, Ruprecht-Karls-Universität Heidelberg, Heidelberg, Germany
- ^{58b}Physikalisches Institut, Ruprecht-Karls-Universität Heidelberg, Heidelberg, Germany
- ^{58c}ZITI Institut für technische Informatik, Ruprecht-Karls-Universität Heidelberg, Mannheim, Germany
- ⁵⁹Faculty of Applied Information Science, Hiroshima Institute of Technology, Hiroshima, Japan
- ⁶⁰Department of Physics, Indiana University, Bloomington, Indiana, USA
- ⁶¹Institut für Astro- und Teilchenphysik, Leopold-Franzens-Universität, Innsbruck, Austria
- ⁶²University of Iowa, Iowa City, Iowa, USA
- ⁶³Department of Physics and Astronomy, Iowa State University, Ames, Iowa, USA
- ⁶⁴Joint Institute for Nuclear Research, JINR Dubna, Dubna, Russia
- ⁶⁵KEK, High Energy Accelerator Research Organization, Tsukuba, Japan
- ⁶⁶Graduate School of Science, Kobe University, Kobe, Japan
- ⁶⁷Faculty of Science, Kyoto University, Kyoto, Japan
- ⁶⁸Kyoto University of Education, Kyoto, Japan

- ⁶⁹*Department of Physics, Kyushu University, Fukuoka, Japan*
- ⁷⁰*Instituto de Física La Plata, Universidad Nacional de La Plata and CONICET, La Plata, Argentina*
- ⁷¹*Physics Department, Lancaster University, Lancaster, United Kingdom*
- ^{72a}*INFN Sezione di Lecce, Italy*
- ^{72b}*Dipartimento di Matematica e Fisica, Università del Salento, Lecce, Italy*
- ⁷³*Oliver Lodge Laboratory, University of Liverpool, Liverpool, United Kingdom*
- ⁷⁴*Department of Physics, Jožef Stefan Institute and University of Ljubljana, Ljubljana, Slovenia*
- ⁷⁵*School of Physics and Astronomy, Queen Mary University of London, London, United Kingdom*
- ⁷⁶*Department of Physics, Royal Holloway University of London, Surrey, United Kingdom*
- ⁷⁷*Department of Physics and Astronomy, University College London, London, United Kingdom*
- ⁷⁸*Louisiana Tech University, Ruston, Louisiana, USA*
- ⁷⁹*Laboratoire de Physique Nucléaire et de Hautes Energies, UPMC and Université Paris-Diderot and CNRS/IN2P3, Paris, France*
- ⁸⁰*Fysiska institutionen, Lunds universitet, Lund, Sweden*
- ⁸¹*Departamento de Física Teórica C-15, Universidad Autónoma de Madrid, Madrid, Spain*
- ⁸²*Institut für Physik, Universität Mainz, Mainz, Germany*
- ⁸³*School of Physics and Astronomy, University of Manchester, Manchester, United Kingdom*
- ⁸⁴*CPPM, Aix-Marseille Université and CNRS/IN2P3, Marseille, France*
- ⁸⁵*Department of Physics, University of Massachusetts, Amherst, Massachusetts, USA*
- ⁸⁶*Department of Physics, McGill University, Montreal, Quebec, Canada*
- ⁸⁷*School of Physics, University of Melbourne, Victoria, Australia*
- ⁸⁸*Department of Physics, The University of Michigan, Ann Arbor, Michigan, USA*
- ⁸⁹*Department of Physics and Astronomy, Michigan State University, East Lansing, Michigan, USA*
- ^{90a}*INFN Sezione di Milano, Italy*
- ^{90b}*Dipartimento di Fisica, Università di Milano, Milano, Italy*
- ⁹¹*B.I. Stepanov Institute of Physics, National Academy of Sciences of Belarus, Minsk, Republic of Belarus*
- ⁹²*National Scientific and Educational Centre for Particle and High Energy Physics, Minsk, Republic of Belarus*
- ⁹³*Department of Physics, Massachusetts Institute of Technology, Cambridge, Massachusetts, USA*
- ⁹⁴*Group of Particle Physics, University of Montreal, Montreal, Quebec, Canada*
- ⁹⁵*P.N. Lebedev Institute of Physics, Academy of Sciences, Moscow, Russia*
- ⁹⁶*Institute for Theoretical and Experimental Physics (ITEP), Moscow, Russia*
- ⁹⁷*Moscow Engineering and Physics Institute (MEPhI), Moscow, Russia*
- ⁹⁸*D.V.Skobel'tsyn Institute of Nuclear Physics, M.V.Lomonosov Moscow State University, Moscow, Russia*
- ⁹⁹*Fakultät für Physik, Ludwig-Maximilians-Universität München, München, Germany*
- ¹⁰⁰*Max-Planck-Institut für Physik (Werner-Heisenberg-Institut), München, Germany*
- ¹⁰¹*Nagasaki Institute of Applied Science, Nagasaki, Japan*
- ¹⁰²*Graduate School of Science and Kobayashi-Maskawa Institute, Nagoya University, Nagoya, Japan*
- ^{103a}*INFN Sezione di Napoli, Italy*
- ^{103b}*Dipartimento di Scienze Fisiche, Università di Napoli, Napoli, Italy*
- ¹⁰⁴*Department of Physics and Astronomy, University of New Mexico, Albuquerque, New Mexico, USA*
- ¹⁰⁵*Institute for Mathematics, Astrophysics and Particle Physics, Radboud University Nijmegen/Nikhef, Nijmegen, Netherlands*
- ¹⁰⁶*Nikhef National Institute for Subatomic Physics and University of Amsterdam, Amsterdam, Netherlands*
- ¹⁰⁷*Department of Physics, Northern Illinois University, DeKalb, Illinois, USA*
- ¹⁰⁸*Budker Institute of Nuclear Physics, SB RAS, Novosibirsk, Russia*
- ¹⁰⁹*Department of Physics, New York University, New York, New York, USA*
- ¹¹⁰*Ohio State University, Columbus, Ohio, USA*
- ¹¹¹*Faculty of Science, Okayama University, Okayama, Japan*
- ¹¹²*Homer L. Dodge Department of Physics and Astronomy, University of Oklahoma, Norman, Oklahoma, USA*
- ¹¹³*Department of Physics, Oklahoma State University, Stillwater, Oklahoma, USA*
- ¹¹⁴*Palacký University, RCPTM, Olomouc, Czech Republic*
- ¹¹⁵*Center for High Energy Physics, University of Oregon, Eugene, Oregon, USA*
- ¹¹⁶*LAL, Université Paris-Sud and CNRS/IN2P3, Orsay, France*
- ¹¹⁷*Graduate School of Science, Osaka University, Osaka, Japan*
- ¹¹⁸*Department of Physics, University of Oslo, Oslo, Norway*
- ¹¹⁹*Department of Physics, Oxford University, Oxford, United Kingdom*
- ^{120a}*INFN Sezione di Pavia, Italy*
- ^{120b}*Dipartimento di Fisica, Università di Pavia, Pavia, Italy*
- ¹²¹*Department of Physics, University of Pennsylvania, Philadelphia, Pennsylvania, USA*
- ¹²²*Petersburg Nuclear Physics Institute, Gatchina, Russia*
- ^{123a}*INFN Sezione di Pisa, Italy*
- ^{123b}*Dipartimento di Fisica E. Fermi, Università di Pisa, Pisa, Italy*
- ¹²⁴*Department of Physics and Astronomy, University of Pittsburgh, Pittsburgh, Pennsylvania, USA*

- ^{125a}Laboratorio de Instrumentacao e Fisica Experimental de Particulas-LIP, Lisboa, Portugal
- ^{125b}Departamento de Fisica Teorica y del Cosmos and CAFPE, Universidad de Granada, Granada, Spain
- ¹²⁶Institute of Physics, Academy of Sciences of the Czech Republic, Praha, Czech Republic
- ¹²⁷Czech Technical University in Prague, Praha, Czech Republic
- ¹²⁸Faculty of Mathematics and Physics, Charles University in Prague, Praha, Czech Republic
- ¹²⁹State Research Center Institute for High Energy Physics, Protvino, Russia
- ¹³⁰Particle Physics Department, Rutherford Appleton Laboratory, Didcot, United Kingdom
- ¹³¹Physics Department, University of Regina, Regina, Saskatchewan, Canada
- ¹³²Ritsumeikan University, Kusatsu, Shiga, Japan
- ^{133a}INFN Sezione di Roma I, Italy
- ^{133b}Dipartimento di Fisica, Università La Sapienza, Roma, Italy
- ^{134a}INFN Sezione di Roma Tor Vergata, Italy
- ^{134b}Dipartimento di Fisica, Università di Roma Tor Vergata, Roma, Italy
- ^{135a}INFN Sezione di Roma Tre, Italy
- ^{135b}Dipartimento di Matematica e Fisica, Università Roma Tre, Roma, Italy
- ^{136a}Faculté des Sciences Ain Chock, Réseau Universitaire de Physique des Hautes Energies-Université Hassan II, Casablanca, Morocco
- ^{136b}Centre National de l'Energie des Sciences Techniques Nucleaires, Rabat, Morocco
- ^{136c}Faculté des Sciences Semlalia, Université Cadi Ayyad, LPHEA-Marrakech, Morocco
- ^{136d}Faculté des Sciences, Université Mohamed Premier and LPTPM, Oujda, Morocco
- ^{136e}Faculté des sciences, Université Mohammed V-Agdal, Rabat, Morocco
- ¹³⁷DSM/IRFU (Institut de Recherches sur les Lois Fondamentales de l'Univers), CEA Saclay (Commissariat à l'Energie Atomique et aux Energies Alternatives), Gif-sur-Yvette, France
- ¹³⁸Santa Cruz Institute for Particle Physics, University of California Santa Cruz, Santa Cruz, California, USA
- ¹³⁹Department of Physics, University of Washington, Seattle, Washington, USA
- ¹⁴⁰Department of Physics and Astronomy, University of Sheffield, Sheffield, United Kingdom
- ¹⁴¹Department of Physics, Shinshu University, Nagano, Japan
- ¹⁴²Fachbereich Physik, Universität Siegen, Siegen, Germany
- ¹⁴³Department of Physics, Simon Fraser University, Burnaby, British Columbia, Canada
- ¹⁴⁴SLAC National Accelerator Laboratory, Stanford, California, USA
- ^{145a}Faculty of Mathematics, Physics & Informatics, Comenius University, Bratislava, Slovak Republic
- ^{145b}Department of Subnuclear Physics, Institute of Experimental Physics of the Slovak Academy of Sciences, Kosice, Slovak Republic
- ^{146a}Department of Physics, University of Cape Town, Cape Town, South Africa
- ^{146b}Department of Physics, University of Johannesburg, Johannesburg, South Africa
- ^{146c}School of Physics, University of the Witwatersrand, Johannesburg, South Africa
- ^{147a}Department of Physics, Stockholm University, Sweden
- ^{147b}The Oskar Klein Centre, Stockholm, Sweden
- ¹⁴⁸Physics Department, Royal Institute of Technology, Stockholm, Sweden
- ¹⁴⁹Departments of Physics & Astronomy and Chemistry, Stony Brook University, Stony Brook, New York, USA
- ¹⁵⁰Department of Physics and Astronomy, University of Sussex, Brighton, United Kingdom
- ¹⁵¹School of Physics, University of Sydney, Sydney, Australia
- ¹⁵²Institute of Physics, Academia Sinica, Taipei, Taiwan
- ¹⁵³Department of Physics, Technion: Israel Institute of Technology, Haifa, Israel
- ¹⁵⁴Raymond and Beverly Sackler School of Physics and Astronomy, Tel Aviv University, Tel Aviv, Israel
- ¹⁵⁵Department of Physics, Aristotle University of Thessaloniki, Thessaloniki, Greece
- ¹⁵⁶International Center for Elementary Particle Physics and Department of Physics, The University of Tokyo, Tokyo, Japan
- ¹⁵⁷Graduate School of Science and Technology, Tokyo Metropolitan University, Tokyo, Japan
- ¹⁵⁸Department of Physics, Tokyo Institute of Technology, Tokyo, Japan
- ¹⁵⁹Department of Physics, University of Toronto, Toronto, Ontario, Canada
- ^{160a}TRIUMF, Vancouver, British Columbia, Canada
- ^{160b}Department of Physics and Astronomy, York University, Toronto, Ontario, Canada
- ¹⁶¹Faculty of Pure and Applied Sciences, University of Tsukuba, Tsukuba, Japan
- ¹⁶²Department of Physics and Astronomy, Tufts University, Medford, Massachusetts, USA
- ¹⁶³Centro de Investigaciones, Universidad Antonio Narino, Bogota, Colombia
- ¹⁶⁴Department of Physics and Astronomy, University of California Irvine, Irvine, California, USA
- ^{165a}INFN Gruppo Collegato di Udine, Italy
- ^{165b}ICTP, Trieste, Italy
- ^{165c}Dipartimento di Chimica, Fisica e Ambiente, Università di Udine, Udine, Italy
- ¹⁶⁶Department of Physics, University of Illinois, Urbana, Illinois, USA
- ¹⁶⁷Department of Physics and Astronomy, University of Uppsala, Uppsala, Sweden

- ¹⁶⁸*Instituto de Física Corpuscular (IFIC) and Departamento de Física Atómica, Molecular y Nuclear and Departamento de Ingeniería Electrónica and Instituto de Microelectrónica de Barcelona (IMB-CNM), University of Valencia and CSIC, Valencia, Spain*
- ¹⁶⁹*Department of Physics, University of British Columbia, Vancouver, British Columbia, Canada*
- ¹⁷⁰*Department of Physics and Astronomy, University of Victoria, Victoria, British Columbia, Canada*
- ¹⁷¹*Department of Physics, University of Warwick, Coventry, United Kingdom*
- ¹⁷²*Waseda University, Tokyo, Japan*
- ¹⁷³*Department of Particle Physics, The Weizmann Institute of Science, Rehovot, Israel*
- ¹⁷⁴*Department of Physics, University of Wisconsin, Madison, Wisconsin, USA*
- ¹⁷⁵*Fakultät für Physik und Astronomie, Julius-Maximilians-Universität, Würzburg, Germany*
- ¹⁷⁶*Fachbereich C Physik, Bergische Universität Wuppertal, Wuppertal, Germany*
- ¹⁷⁷*Department of Physics, Yale University, New Haven, Connecticut, USA*
- ¹⁷⁸*Yerevan Physics Institute, Yerevan, Armenia*
- ¹⁷⁹*Centre de Calcul de l'Institut National de Physique Nucléaire et de Physique des Particules (IN2P3), Villeurbanne, France*

^aDeceased.

^bAlso at Department of Physics, King's College London, London, United Kingdom.

^cAlso at Laboratório de Instrumentação e Física Experimental de Partículas-LIP, Lisboa, Portugal.

^dAlso at Institute of Physics, Azerbaijan Academy of Sciences, Baku, Azerbaijan.

^eAlso at Faculdade de Ciências and CFNUL, Universidade de Lisboa, Lisboa, Portugal.

^fAlso at Particle Physics Department, Rutherford Appleton Laboratory, Didcot, United Kingdom.

^gAlso at TRIUMF, Vancouver, British Columbia, Canada.

^hAlso at Department of Physics, California State University, Fresno, CA, USA.

ⁱAlso at Novosibirsk State University, Novosibirsk, Russia.

^jAlso at Department of Physics, University of Coimbra, Coimbra, Portugal.

^kAlso at Università di Napoli Parthenope, Napoli, Italy.

^lAlso at Institute of Particle Physics (IPP), Canada.

^mAlso at Department of Physics, Middle East Technical University, Ankara, Turkey.

ⁿAlso at Louisiana Tech University, Ruston, LA, USA.

^oAlso at Dep Física and CEFITEC of Faculdade de Ciências e Tecnologia, Universidade Nova de Lisboa, Caparica, Portugal.

^pAlso at CPPM, Aix-Marseille Université and CNRS/IN2P3, Marseille, France.

^qAlso at Department of Physics and Astronomy, Michigan State University, East Lansing, MI, USA.

^rAlso at Department of Financial and Management Engineering, University of the Aegean, Chios, Greece.

^sAlso at Institutio Catalana de Recerca i Estudis Avancats, ICREA, Barcelona, Spain.

^tAlso at Department of Physics, University of Cape Town, Cape Town, South Africa.

^uAlso at CERN, Geneva, Switzerland.

^vAlso at O Chadai Academic Production, Ochanomizu University, Tokyo, Japan.

^wAlso at Manhattan College, New York, NY, USA.

^xAlso at Institute of Physics, Academia Sinica, Taipei, Taiwan.

^yAlso at School of Physics and Engineering, Sun Yat-sen University, Guanzhou, China.

^zAlso at Academia Sinica Grid Computing, Institute of Physics, Academia Sinica, Taipei, Taiwan.

^{aa}Also at Laboratoire de Physique Nucléaire et de Hautes Energies, UPMC and Université Paris-Diderot and CNRS/IN2P3, Paris, France.

^{bb}Also at School of Physical Sciences, National Institute of Science Education and Research, Bhubaneswar, India.

^{cc}Also at Dipartimento di Fisica, Università La Sapienza, Roma, Italy.

^{dd}Also at DSM/IRFU (Institut de Recherches sur les Lois Fondamentales de l'Univers), CEA Saclay (Commissariat à l'Énergie Atomique et aux Énergies Alternatives), Gif-sur-Yvette, France.

^{ee}Also at Moscow Institute of Physics and Technology State University, Dolgoprudny, Russia.

^{ff}Also at Section de Physique, Université de Genève, Geneva, Switzerland.

^{gg}Also at Departamento de Física, Universidade de Minho, Braga, Portugal.

^{hh}Also at Department of Physics, The University of Texas at Austin, Austin, TX, USA.

ⁱⁱAlso at Institute for Particle and Nuclear Physics, Wigner Research Centre for Physics, Budapest, Hungary.

^{jj}Also at DESY, Hamburg and Zeuthen, Germany.

^{kk}Also at International School for Advanced Studies (SISSA), Trieste, Italy.

^{ll}Also at Department of Physics and Astronomy, University of South Carolina, Columbia, SC, USA.

^{mm}Also at Faculty of Physics, M.V.Lomonosov Moscow State University, Moscow, Russia.

ⁿⁿAlso at Nevis Laboratory, Columbia University, Irvington, NY, USA.

^{oo}Also at Physics Department, Brookhaven National Laboratory, Upton, NY, USA.

^{pp}Also at Moscow Engineering and Physics Institute (MEPhI), Moscow, Russia.

^{qq}Also at Department of Physics, Oxford University, Oxford, United Kingdom.

^{rr}Also at Institut für Experimentalphysik, Universität Hamburg, Hamburg, Germany.

^{ss}Also at Department of Physics, The University of Michigan, Ann Arbor, MI, USA.

^{tt}Also at Discipline of Physics, University of KwaZulu-Natal, Durban, South Africa.

^{uu}atlas.publication@cern.ch